

Search for dark matter at $\sqrt{s} = 13$ TeV in final states containing an energetic photon and large missing transverse momentum with the ATLAS detector

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Abstract Results of a search for physics beyond the Standard Model in events containing an energetic photon and large missing transverse momentum with the ATLAS detector at the Large Hadron Collider are reported. As the number of events observed in data, corresponding to an integrated luminosity of 36.1 fb^{-1} of proton–proton collisions at a centre-of-mass energy of 13 TeV, is in agreement with the Standard Model expectations, model-independent limits are set on the fiducial cross section for the production of events in this final state. Exclusion limits are also placed in models where dark-matter candidates are pair-produced. For dark-matter production via an axial-vector or a vector mediator in the s -channel, this search excludes mediator masses below 750–1200 GeV for dark-matter candidate masses below 230–480 GeV at 95% confidence level, depending on the couplings. In an effective theory of dark-matter production, the limits restrict the value of the suppression scale M_* to be above 790 GeV at 95% confidence level. A limit is also reported on the production of a high-mass scalar resonance by processes beyond the Standard Model, in which the resonance decays to $Z\gamma$ and the Z boson subsequently decays into neutrinos.

Contents

| | |
|---|---|
| 1 Introduction | 1 |
| 2 The ATLAS detector | 2 |
| 3 Monte Carlo simulation samples | 3 |
| 4 Event reconstruction | 4 |
| 5 Event selection | 5 |
| 6 Background estimation | 6 |
| 6.1 $W\gamma$ and $Z\gamma$ backgrounds | 6 |
| 6.2 γ +jets background | 7 |
| 6.3 Fake photons from misidentified electrons | 7 |
| 6.4 Fake photons from misidentified jets | 7 |

| | |
|--|----|
| 6.5 Final background estimation | 7 |
| 6.5.1 Background-only inclusive-SR fit | 8 |
| 6.5.2 Background-only multiple-bin fit | 8 |
| 7 Results | 8 |
| 8 Systematic uncertainties | 10 |
| 9 Interpretation of results | 11 |
| 10 Conclusion | 15 |
| References | 15 |

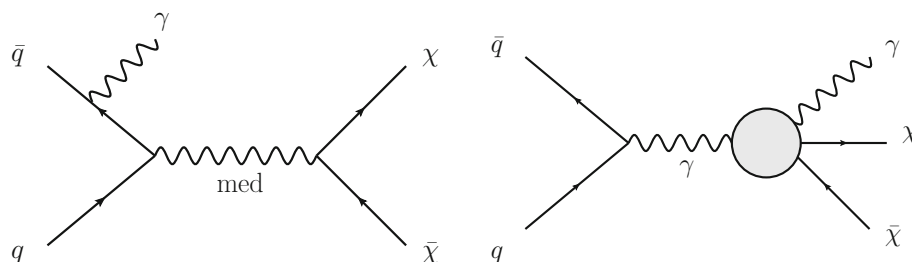
1 Introduction

Multiple theories of physics beyond the Standard Model (BSM) predict a high production rate of events containing a photon with a high transverse energy (E_T^γ) and large missing transverse momentum (E_T^{miss} , with magnitude E_T^{miss}) referred to as $\gamma + E_T^{\text{miss}}$ events, in pp collisions. The search for energetic $\gamma + E_T^{\text{miss}}$ events, whose rates have a low expected contribution from Standard Model (SM) processes, can thus provide sensitivity to new physics models [1–5], also related to dark matter (DM). Although the existence of DM is well established [6], its nature is yet unknown. One candidate is a weakly interacting massive particle (WIMP, also denoted by χ) that interacts with SM particles with a strength similar to the weak interaction. If WIMPs interact with quarks via a mediator particle, pairs of WIMPs are produced in pp collisions at sufficiently high energy. The $\chi\bar{\chi}$ pair is invisible to the detector, but the radiation of an initial-state photon in $q\bar{q} \rightarrow \chi\bar{\chi}$ interactions [7] can produce detectable $\gamma + E_T^{\text{miss}}$ events.

Effective field theories (EFT) with various forms of interaction between the WIMPs and the SM particles are a powerful model-independent approach for the interpretation of DM production in pp collisions [7]. However, the typical momentum transfer in pp collisions at the LHC can often exceed the cut-off scale below which the EFT approximation is valid. Simplified models that involve the explicit pro-

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Fig. 1 Pair production of dark-matter particles ($\chi\bar{\chi}$) in association with a photon via an explicit s -channel mediator (left), or via an effective $\gamma\gamma\chi\bar{\chi}$ vertex (right)



duction of the intermediate state, as shown in Fig. 1 (left), avoid this limitation. This paper focuses on simplified models assuming Dirac-fermion DM candidates produced via an s -channel mediator with vector or axial-vector interactions [8–10]. There are five free parameters in this model: the WIMP mass m_χ , the mediator mass m_{med} , the width of the mediator Γ_{med} , the coupling g_q of the mediator to quarks, and the coupling g_χ of the mediator to the dark-matter particle. In the limit of a large mediator mass, these simplified models map onto the EFT operators, with the suppression scale¹ M_* linked to m_{med} by the relation $M_* = m_{\text{med}}/\sqrt{g_q g_\chi}$ [11].

The paper also considers a specific dimension-7 EFT operator with direct couplings between DM and electroweak (EW) bosons, for which there is neither a corresponding simplified model nor a simplified model yielding similar kinematic distributions implemented in an event generator [10, 12]. The process describing a contact interaction of type $\gamma\gamma\chi\bar{\chi}$ is shown in Fig. 1 (right). In this model, DM production proceeds via $q\bar{q} \rightarrow \gamma \rightarrow \gamma\chi\bar{\chi}$, generating an energetic photon without requiring initial-state radiation. There are four free parameters in this model: the EW coupling strengths k_1 and k_2 (which respectively control the strength of the coupling to the SM U(1) and SU(2) gauge sectors), m_χ , and the suppression scale M_* .

Many BSM models [13, 14] introduce new bosons through either an extension of the Higgs sector or additional gauge fields. In some of those, the bosons are predicted to decay into electroweak gauge bosons: the analysis presented here also searches for such a resonance decaying into $Z\gamma$, which would lead to an excess of energetic $\gamma + E_{\text{T}}^{\text{miss}}$ events when the Z boson subsequently decays to neutrinos.

The ATLAS and CMS collaborations have reported limits in various models based on searches for an excess of $\gamma + E_{\text{T}}^{\text{miss}}$ events using pp collisions at centre-of-mass energies of $\sqrt{s} = 7$ and 8 TeV (LHC Run 1) and with the first LHC Run-2 data collected in 2015 at a centre-of-mass energy of 13 TeV [15–19]. A $\chi\bar{\chi}$ pair can also be produced in association with other objects leading to different $X + E_{\text{T}}^{\text{miss}}$ signatures, where X can be a jet, a W boson, a Z boson or a Higgs boson. DM searches are hence performed in a vari-

ety of complementary final states [20–24]. The $\gamma + E_{\text{T}}^{\text{miss}}$ final state has the advantage of a clean signature providing a good complementarity with respect to the other $X + E_{\text{T}}^{\text{miss}}$ processes. Moreover it also offers the unique possibility to probe for DM models in which the photon does not come from initial-state radiation. This paper reports the results of a search for dark matter and for a BSM $Z\gamma$ resonance in $\gamma + E_{\text{T}}^{\text{miss}}$ events in pp collisions at a centre-of-mass energy of $\sqrt{s} = 13$ TeV using the Run-2 data collected in 2015 and 2016, corresponding to an integrated luminosity of 36.1 fb^{-1} . As described in Sect. 5, this search follows a strategy similar to that implemented in Ref. [17], but with multiple signal regions optimised to take advantage of the tenfold increase in integrated luminosity.

The paper is organised as follows. A brief description of the ATLAS detector is given in Sect. 2. The signal and background Monte Carlo (MC) simulation samples used are described in Sect. 3. The reconstruction of physics objects is explained in Sect. 4, and the event selection is described in Sect. 5. Estimation of the SM backgrounds is outlined in Sect. 6. The results are described in Sect. 7 and the systematic uncertainties are given in Sect. 8. The interpretation of results in terms of models of pair production of dark-matter candidates and of BSM production of a high-mass $Z\gamma$ resonance is described in Sect. 9. A summary is given in Sect. 10.

2 The ATLAS detector

The ATLAS detector [25] is a multipurpose particle physics apparatus with a forward–backward symmetric cylindrical geometry and near 4π coverage in solid angle.² The inner tracking detector (ID), covering the pseudorapidity range $|\eta| < 2.5$, consists of a silicon pixel detector including the insertable B-layer [26, 27], which was added around a new, smaller-radius beam-pipe before the start of Run 2; a silicon microstrip detector; and, for $|\eta| < 2.0$, a straw-tube transi-

¹ The suppression scale, also referred to as Λ , is the effective mass scale of particles that are integrated out in an EFT. The non-renormalisable operators are suppressed by powers of $1/M_*$.

² ATLAS uses a right-handed coordinate system with its origin at the nominal interaction point (IP) in the centre of the detector and the z -axis along the beam pipe. The x -axis points from the IP to the centre of the LHC ring, and the y -axis points upward. Cylindrical coordinates (r, ϕ) are used in the transverse plane, ϕ being the azimuthal angle around the beam pipe. The pseudorapidity is defined in terms of the polar θ angle as $\eta = -\ln[\tan(\theta/2)]$.

tion radiation tracker (TRT). The ID is surrounded by a thin superconducting solenoid which provides a 2 T magnetic field. A high-granularity lead/liquid-argon sampling electromagnetic calorimeter (EM) covers the region $|\eta| < 3.2$. It is segmented longitudinally in shower depth. The first layer has a high granularity in the η direction in order to provide an efficient discrimination between single-photon showers and two overlapping photons originating from a π^0 decay. The second layer is where most of the energy, deposited in the calorimeter by electron- or photon-initiated electromagnetic showers, is collected. Significant energy deposits can be left in the third layer by very high energy showers; this layer can also be used to correct for energy leakage beyond the electromagnetic calorimeter. A steel/scintillator-tile hadronic calorimeter covers the range $|\eta| < 1.7$, while the liquid-argon technology with either copper or tungsten as the absorber material is used for the hadronic calorimeters in the end-cap region $1.5 < |\eta| < 3.2$ and for electromagnetic and hadronic measurements in the forward region up to $|\eta| = 4.9$. A muon spectrometer (MS) surrounds the calorimeters. It consists of three large air-core superconducting toroidal magnet systems, precision tracking chambers providing accurate muon tracking out to $|\eta| = 2.7$, and fast detectors for triggering in the region $|\eta| < 2.4$. A two-level trigger system is used to select events for offline analysis [28].

3 Monte Carlo simulation samples

Several simulated MC samples are used to estimate the signal acceptance, the detector efficiency and various SM background contributions. For all the DM samples considered here, the values of the free parameters were chosen following the recommendations given in Ref. [10].

Samples of DM production in simplified models are generated via an s -channel mediator with axial-vector interactions. The program MG5_aMC@NLO v2.4.3 [29] is used in conjunction with PYTHIA v8.212 [30] with the parameter values set according to the ATLAS tune A14 [31]. The parton distribution function (PDF) set used is NNPDF3.0 at next-to-leading order (NLO) [32] with $\alpha_s = 0.118$. The g_q coupling is set to be universal in quark flavour and equal to 0.25, the g_χ coupling is set to 1.0, and Γ_{med} is computed as the minimum width allowed given the couplings and masses. As shown in Ref. [10], $\Gamma_{\text{med}}/m_{\text{med}}$ varies between 2 and 6% for the values probed here. Different choices of the couplings and a model with a vector mediator are also considered, as described in Sect. 9. The generation was updated with respect to the 2015 data analysis [17] by using the DMsimp [33] implementation of the model at NLO. Events are generated with parameters spanning a grid of points in the m_χ - m_{med} plane.

For DM samples corresponding to an EFT model involving dimension-7 operators with a contact interaction of type

$\gamma\gamma\chi\bar{\chi}$, the parameters which only influence the cross section are set to $k_1 = k_2 = 1.0$ and $M_* = 3.0$ TeV [10]. A scan over a range of values of m_χ is performed. Events are generated with MG5_aMC@NLO v2.2.3 and the PDF set NNPDF3.0 at leading order (LO) with $\alpha_s = 0.130$, in conjunction with PYTHIA v8.186, using the ATLAS tune A14.

For DM signal generation in both the simplified and EFT models, a photon with at least $E_T^\gamma = 130$ GeV is required at the matrix-element level in MG5_aMC@NLO.

The samples used in the search for a BSM high-mass scalar resonance decaying to $Z\gamma$ are generated using POWHEG-BOX v1 [34], with the CT10 PDF set [35] and PYTHIA v8.210 for the showering with the AZNLO tune [36] based on the CTEQ6L1 PDF set [37]. The simulated heavy scalar resonance X of very narrow width (4 MeV), with masses in the range 2 to 5 TeV, is produced through gluon-gluon fusion and then assumed to decay exclusively to $Z\gamma$.

For all the signal samples described above, the EvtGen v1.2.0 program [38] is used for properties of the bottom and charm hadron decays.

For $W\gamma$ and $Z\gamma$ backgrounds, events containing a charged lepton (e , μ or τ) and a neutrino, a pair of neutrinos ($\nu\nu$) or a pair of charged leptons ($\ell\ell$) together with a photon and associated jets are simulated using the SHERPA v2.1.1 event generator [39]. The matrix elements including all diagrams with three electroweak couplings are calculated with up to three partons at LO and merged with SHERPA parton showers [40] using the ME+PS@LO prescription [41]. The CT10 PDF set is used in conjunction with a dedicated parton shower tuning developed by the SHERPA authors. For Z events with the Z boson decaying to a $\ell\ell$ pair a requirement on the dilepton invariant mass of $m_{\ell\ell} > 10$ GeV is applied at event generator level.

Events containing a photon with associated jets are also simulated using SHERPA v2.1.1 [39], generated in several bins of E_T^γ with lower edges ranging from 35 GeV to 4 TeV. The matrix elements are calculated at LO with up to three or four partons and merged with SHERPA parton showers using the ME+PS@LO prescription. The CT10 PDF set is used in conjunction with the dedicated parton shower tuning.

For W/Z +jets backgrounds, events containing W or Z bosons with associated jets are simulated using SHERPA v2.2.0. The matrix elements are calculated for up to four partons at LO and two partons at NLO using the Comix [42] and OpenLoops [43] matrix-element generators and merged with SHERPA parton showers using the ME+PS@NLO prescription [44]. The NNPDF3.0 PDF set at next-to-next-to-leading order (NNLO) is used. As in the case of the γ +jets samples, the dedicated parton shower tuning is used. The W/Z +jets events are normalised to the NNLO inclusive cross sections [45].

Table 1 summarises the choices made in the generation of MC samples used in the analysis.

Table 1 Details of the generation of the signal samples and of the SM background processes considered in the analysis

| Process | Event generators used | PDF sets | Order | Requirements |
|-----------------|------------------------------------|---------------------|--------|--------------------------------|
| DMsimp model | MG5_aMC@NLO v2.4.3 + PYTHIA v8.212 | NNPDF30_nlo_as_0118 | NLO | $E_T^\gamma > 130$ GeV |
| EFT model | MG5_aMC@NLO v2.2.3 + PYTHIA v8.186 | NNPDF30_lo_as_0130 | LO | $E_T^\gamma > 130$ GeV |
| BSM resonance | POWHEG-BOX v1 + PYTHIA v8.210 | CT10 | NLO | – |
| $W/Z\gamma$ | SHERPA v2.1.1 | CT10 | LO | For Z: $m_{\ell\ell} > 10$ GeV |
| γ + jets | SHERPA v2.1.1 | CT10 | LO | – |
| W/Z +jets | SHERPA v2.2.0 | NNPDF30_nnlo | LO/NLO | – |

Multiple pp interactions in the same or neighbouring bunch crossings (referred to as pile-up) superimposed on the hard physics process are simulated with the minimum-bias processes of PYTHIA v8.186 using the A2 tune [46] and the MSTW2008LO PDF set [47]. Simulated events are reweighted so that the distribution of the expected number of collisions per bunch crossing, $\langle\mu\rangle$, matches the one observed in data, which has a mean value of 13.7 (24.2) in 2015 (2016) data.

All generated event samples are processed with a full ATLAS detector simulation [48] based on GEANT4 [49]. The simulated events are reconstructed and analysed with the same analysis chain as used for the data, utilising the same trigger and event selection criteria discussed in Sect. 5.

4 Event reconstruction

Photons are reconstructed from clusters of energy deposits in the electromagnetic calorimeter measured in projective towers. Clusters without matching tracks are classified as unconverted photon candidates. A photon candidate containing clusters that can be matched to tracks is considered as a converted photon candidate [50]. The photon energy is corrected by applying the energy scales measured with $Z \rightarrow e^+e^-$ decays [51]. The trajectory of the photon is reconstructed using the longitudinal (shower depth) segmentation of the calorimeters and a constraint from the average collision point of the proton beams. For converted photons, the position of the conversion vertex is also used if tracks from the conversion have hits in the silicon detectors. Identification requirements are applied in order to reduce the contamination from π^0 or other neutral hadrons decaying to two photons. The photon identification is based on the profile of the energy deposits in the first and second layers of the electromagnetic calorimeter. Candidate photons are required to have $E_T^\gamma > 10$ GeV, to satisfy the “loose” identification criteria defined in Ref. [52] and to be within $|\eta| < 2.37$. Photons used in the event selection must additionally satisfy the “tight” identification criteria [52], have $|\eta| < 1.37$ or $1.52 < |\eta| < 2.37$ and be isolated by

requiring the energy in the calorimeters in a cone of size $\Delta R = \sqrt{(\Delta\eta)^2 + (\Delta\phi)^2} = 0.4$ around the cluster barycentre, excluding the energy associated with the photon cluster, to be less than 2.45 GeV + $0.022 \times E_T^\gamma$. This cone energy is corrected for the leakage of the photon energy from the central core and for the effects of pile-up [51]. In addition, the scalar sum of the p_T of non-conversion tracks in a cone of size $\Delta R = 0.2$ around the cluster barycentre is required to be less than $0.05 \times E_T^\gamma$.

Electrons are reconstructed from clusters in the electromagnetic calorimeter which are matched to a track in the ID. The criteria for their identification, and the calibration steps, are similar to those used for photons. Electron candidates must fulfil the “medium” identification requirement of Ref. [51]. Muons are identified either as a combined track in the MS and ID systems, or as an ID track that, once extrapolated to the MS, is associated with at least one track segment in the MS. Extrapolated muons are also considered; they are reconstructed from an MS track which is required to be compatible with originating from the nominal interaction point. Muon candidates must pass the “medium” identification requirement [53]. The significance of the transverse impact parameter, defined as the transverse impact parameter d_0 divided by its estimated uncertainty, σ_{d_0} , of tracks with respect to the beam line is required to satisfy $|d_0|/\sigma_{d_0} < 5.0$ for electrons and $|d_0|/\sigma_{d_0} < 3.0$ for muons. The longitudinal impact parameter z_0 must satisfy $|z_0 \sin\theta| < 0.5$ mm for both the electrons and muons. Electrons are required to have $p_T > 7$ GeV and $|\eta| < 2.47$, while muons are required to have $p_T > 6$ GeV and $|\eta| < 2.7$.

Jets are reconstructed with the anti- k_t algorithm [54] with a radius parameter $R = 0.4$ from clusters of energy deposits at the electromagnetic scale in the calorimeters [55]. A correction used to calibrate the jet energy to the scale of its constituent particles [56,57] is then applied. Jets are also corrected for contributions from pile-up interactions and a residual correction derived in situ is applied as the last step to jets reconstructed in data [56]. Candidate jets are required to have $p_T > 20$ GeV. In order to suppress pile-up jets, which are mainly at low p_T , a jet vertex tagger [58], based on tracking and vertexing information, is applied for jets with

$p_T < 60$ GeV and $|\eta| < 2.4$. Jets used in the event selection are required to have $p_T > 30$ GeV and $|\eta| < 4.5$. The τ leptons decaying to hadrons and ν_τ are considered as jets as in previous analyses [16, 17].

The removal of overlapping candidate objects is performed in the following order. If any selected electron shares its ID track with a selected muon, the electron is removed and the muon is kept, in order to remove electron candidates originating from muon bremsstrahlung followed by photon conversion. If an electron lies a distance $\Delta R < 0.2$ of a candidate jet, the jet is removed from the event, while if an electron lies a distance $0.2 < \Delta R < 0.4$ of a jet, the electron is removed. Muons lying a distance $\Delta R < 0.4$ with respect to the remaining candidate jets are removed, except if the number of tracks with $p_T > 0.5$ GeV associated with the jet is less than three. In the latter case, the muon is kept and the jet is discarded. Finally, if a jet lies a distance $\Delta R < 0.4$ of a candidate photon, the jet is removed.

The missing transverse momentum vector $\mathbf{E}_T^{\text{miss}}$ is obtained from the negative vector sum of the momenta of the candidate physics objects, selected as described above. Calorimeter energy deposits and tracks are matched with candidate high- p_T objects in a specific order: electrons with $p_T > 7$ GeV, photons with $E_T^\gamma > 10$ GeV, muons with $p_T > 6$ GeV and jets with $p_T > 20$ GeV [59]. Tracks from the primary vertex³ not associated with any such objects are also taken into account in the $\mathbf{E}_T^{\text{miss}}$ reconstruction (“soft term”) [61].

Corrections are applied to the objects in the simulated samples to account for differences compared to data in object reconstruction, identification and isolation efficiencies for both the leptons and photons. For the photons, the efficiency corrections depend on whether or not they are converted, and on their E_T^γ and η ; for the photons used in this analysis they are generally of the order of 1%.

5 Event selection

The data were collected in pp collisions at $\sqrt{s} = 13$ TeV during 2015 and 2016. The events for the analysis were recorded using a trigger requiring at least one photon candidate above a E_T^γ threshold of 140 GeV to pass “loose” identification requirements, which are based on the shower shapes in the EM calorimeter as well as on the energy leaking into the hadronic calorimeter [62].

For events in the signal regions defined below, the efficiency of the trigger is more than 98.5%. The 1% difference

in the efficiency between data and MC simulation is treated as a systematic uncertainty. Only data satisfying beam, detector and data-quality criteria are considered. The data used for the analysis correspond to an integrated luminosity of 36.1 fb^{-1} . The uncertainty in the integrated luminosity is $\pm 3.2\%$. It is derived following a methodology similar to that detailed in Ref. [63], from a preliminary calibration of the luminosity scale using x - y beam-separation scans performed in August 2015 and May 2016.

Events are removed if they contain a bad-quality photon or jet [50, 64], arising from instrumental problems or non-collision background. Events are required to have a reconstructed primary vertex, as defined in Sect. 4.

Events in the signal regions (SRs) are required to have the leading photon satisfying the criteria defined in Sect. 4 and having $E_T^\gamma > 150$ GeV, which is well above the thresholds used for the MC event generation and for the data-collection trigger. The $|z|$ position, defined as the longitudinal separation between the beamspot position and the intersection of the extrapolated photon trajectory with the beam-line, is required to be smaller than 0.25 m. This criterion provides a powerful rejection against the muons from beam background [17], which can leave significant energy deposits in the calorimeters and hence lead to reconstructed fake photons that do not point back to the primary vertex. It is required that the photon and $\mathbf{E}_T^{\text{miss}}$ do not overlap in the azimuthal plane: $\Delta\phi(\gamma, \mathbf{E}_T^{\text{miss}}) > 0.4$. To further suppress the background events where the $\mathbf{E}_T^{\text{miss}}$ is fake, e.g. arising from poorly reconstructed objects, a requirement on the ratio $E_T^{\text{miss}}/\sqrt{\Sigma E_T} > 8.5 \text{ GeV}^{1/2}$ is added,⁴ where ΣE_T is calculated as the scalar sum of all p_T from the objects and the tracks contributing to the $\mathbf{E}_T^{\text{miss}}$ reconstruction described in Sect. 4. This requirement mainly rejects the γ +jets background events. Events with more than one jet or with a jet with $\Delta\phi(\text{jets}, \mathbf{E}_T^{\text{miss}}) < 0.4$ are rejected (jet veto), the latter to remove events where $\mathbf{E}_T^{\text{miss}}$ originates from jet mismeasurement. The remaining events with one jet are retained to increase the signal acceptance and reduce systematic uncertainties related to the modelling of initial-state radiation. Events are required to have no electrons or muons passing the requirements for e/μ candidates described in Sect. 4. This lepton veto mainly rejects W/Z events with charged leptons in the final state.

As the production of a pair of dark-matter candidates or of a high-mass BSM $Z(\rightarrow \nu\nu)\gamma$ resonance are both expected to lead to events with large E_T^{miss} , five SRs are defined with different E_T^{miss} requirements: three inclusive (SRI1, SRI2 and SRI3) with increasing E_T^{miss} thresholds and two exclusive (SRE1 and SRE2) with E_T^{miss} in two different ranges. Table 2

³ The primary vertex is defined as the vertex with the highest sum of the squared transverse momenta of its associated tracks. It is reconstructed from at least two associated good-quality tracks with $p_T > 0.4$ GeV [60].

⁴ This ratio simulates the E_T^{miss} significance because the E_T^{miss} resolution due to purely calorimetric measurements scales approximately as $\sqrt{\Sigma E_T}$.

Table 2 Criteria for selecting events in the SRs and the numbers of events selected in data

| Event cleaning | Quality and primary vertex | | | | |
|---------------------------------------|---|------|------|---------|---------|
| Leading photon | $E_T^\gamma > 150 \text{ GeV}$, $ \eta < 1.37$ or $1.52 < \eta < 2.37$, tight, isolated, $ z < 0.25 \text{ m}$, $\Delta\phi(\gamma, \mathbf{E}_T^{\text{miss}}) > 0.4$ | | | | |
| $E_T^{\text{miss}}/\sqrt{\Sigma E_T}$ | $> 8.5 \text{ GeV}^{1/2}$ | | | | |
| Jets | 0 or 1 with $p_T > 30 \text{ GeV}$, $ \eta < 4.5$ and $\Delta\phi(\text{jets}, \mathbf{E}_T^{\text{miss}}) > 0.4$ | | | | |
| Lepton | Veto on e and μ | | | | |
| | SRI1 | SRI2 | SRI3 | SRE1 | SRE2 |
| E_T^{miss} [GeV] | >150 | >225 | >300 | 150–225 | 225–300 |
| Selected events in data | 2400 | 729 | 236 | 1671 | 493 |
| Events with 0 jets | 1559 | 379 | 116 | 1180 | 263 |

shows the criteria for selecting events in the SRs and the number of events selected in data. The fraction of events in which the selected photon is unconverted is about 70% for all regions. The fraction of selected events with no jets increases in the regions with lower E_T^{miss} thresholds and ranges from about 50% to about 70%.

6 Background estimation

The SM background to the $\gamma + E_T^{\text{miss}}$ final state is due to events containing either a true photon or an object misidentified as a photon. The background with a true photon is dominated by the electroweak production of $Z(\rightarrow \nu\nu)\gamma$ events. Secondary contributions come from $W(\rightarrow \ell\nu)\gamma$ and $Z(\rightarrow \ell\ell)\gamma$ production with unidentified electrons, muons or with $\tau \rightarrow \text{hadrons} + \nu_\tau$ decays and from γ +jets events. The contribution from $t\bar{t} + \gamma$ is negligible thanks to the jet veto. The contribution from events where a lepton or a jet is misidentified as a photon arises mainly from W/Z +jets production, with smaller contributions from diboson, multi-jet and top-quark pair production.

All significant background estimates are extrapolated from non-overlapping data samples. A simultaneous fit in background-enriched control regions (CRs) is performed to obtain normalisation factors for the $W\gamma$, $Z\gamma$ and γ +jets backgrounds (see Sects. 6.1 and 6.2), which are then used to estimate backgrounds in the SRs; more details are given in Sects. 6.5.1 and 6.5.2. The same normalisation factor is used for both $Z(\rightarrow \nu\nu)\gamma$ and $Z(\rightarrow \ell\ell)\gamma$ in SR events. The backgrounds due to photons from the misidentification of electrons or jets in processes such as W/Z +jets, diboson and multi-jet events (referred to as fake photons) are estimated using data-driven techniques (see Sects. 6.3 and 6.4).

6.1 $W\gamma$ and $Z\gamma$ backgrounds

For the estimation of the $W/Z\gamma$ background, three control regions are defined for each SR by selecting events with the same criteria used for the various SRs but inverting the lepton vetoes. As the γ +jets background contribution is not significant in these leptonic CRs, the requirement on the ratio $E_T^{\text{miss}}/\sqrt{\Sigma E_T}$ is not applied. In the one-muon control region (1muCR) the $W\gamma$ contribution is enhanced by requiring the presence of a muon; the 1muCR is sufficient to constrain the $W\gamma$ normalisation effectively without the need of a similar one-electron control region, which would be contaminated by γ +jets background. The two-lepton control regions enhance the $Z\gamma$ background by requiring the presence of a pair of muons (2muCR) or electrons (2eleCR). In each case, the CR lepton selection follows the same requirements as the SR lepton veto, with the addition that the leptons must be isolated with “loose” criteria [53] based on information from the calorimeter and tracking systems. In both 1muCR and 2muCR, the E_T^{miss} value is computed disregarding muons, effectively treating them as non-interacting particles, in order to ensure that the E_T^{miss} distributions in those CRs are similar to that in the SR. The same procedure is followed for electrons in 2eleCR. In both the $Z\gamma$ -enriched control regions, the dilepton invariant mass $m_{\ell\ell}$ is required to be greater than 20 GeV, and the invariant mass of the leptons and photon, $m_{\ell\ell\gamma}$, is required to be smaller than 1 TeV in order to avoid probing for potential BSM high-mass $Z\gamma$ resonances. The normalisation of the dominant $Z(\rightarrow \nu\nu)\gamma$ background source is largely constrained by the event yields in 2muCR and 2eleCR. The systematic uncertainty due to the extrapolation of the correction factors from CRs to SRs is taken into account (see Sect. 8).

6.2 γ +jets background

The γ +jets background in the SRs consists of events where the jet is poorly reconstructed and partially lost, creating fake E_T^{miss} . This background, which increased in 2016 data with respect to 2015 data because of the higher pile-up conditions, is suppressed by the large E_T^{miss} and jet- E_T^{miss} azimuthal separation requirements and by the requirement $E_T^{\text{miss}}/\sqrt{\Sigma E_T} > 8.5 \text{ GeV}^{1/2}$ described in Sect. 5. This last requirement reduces the contribution of γ +jets events in SRI1 to less than 10% of the total background, with a negligible effect on the acceptance for signal events. The fraction of γ +jets events decreases with E_T^{miss} and becomes less than 2% of the total background in SRI3. For the estimation of the residual γ +jets background, a specific control region (PhJetCR) enriched in γ +jets events is defined. It uses the same criteria as used for the SRs, but does not apply the requirement on the ratio $E_T^{\text{miss}}/\sqrt{\Sigma E_T}$, and requires $85 \text{ GeV} < E_T^{\text{miss}} < 110 \text{ GeV}$ and azimuthal separation between the photon and E_T^{miss} , $\Delta\phi(\gamma, E_T^{\text{miss}})$, to be smaller than 3.0. The latter selection minimises the contamination from signal events, which is estimated to be at most at the level of 1%. The PhJetCR is the same for all SRs; the systematic uncertainty due to the modelling of the γ +jets background, which affects its extrapolation from the low- E_T^{miss} PhJetCR to the SRs with larger E_T^{miss} , is taken into account (see Sect. 8).

6.3 Fake photons from misidentified electrons

Contributions from processes in which an electron is misidentified as a photon in the SRs are estimated by scaling yields from a sample of $e + E_T^{\text{miss}}$ events by an electron-to-photon misidentification factor. This factor is measured with mutually exclusive samples of e^+e^- and $\gamma + e$ events in data. To establish a pure sample of electrons, the ee and the $e\gamma$ invariant masses (m_{ee} and $m_{e\gamma}$) are both required to be consistent with the Z boson mass to within 10 GeV, and the E_T^{miss} is required to be smaller than 40 GeV. Furthermore, the sidebands to the Z boson mass window are used to estimate and subtract possible contamination from misidentified jets in this sample. The misidentification factor, calculated as the ratio of the number of $\gamma + e$ to the number of e^+e^- events, is parameterised as a function of p_T and pseudorapidity and it varies between 0.59 and 2.5%. Systematic uncertainties in the misidentification factors are evaluated by varying the sideband definition, comparing the results of the method (with or without using the sideband subtraction) with generator-level information about $Z(\rightarrow ee)$ MC events, and comparing the misidentification factors in $Z(\rightarrow ee)$ and $W(\rightarrow e\nu)$ MC events. Background estimates are then also made for the four control regions, 1muCR, 2muCR, 2eleCR and PhJetCR, by applying the electron-to-photon misidenti-

fication factor to events selected with the same criteria used for these regions but requiring an electron instead of a photon. The estimated contribution from this background in the SRs and the associated uncertainty are reported in Sect. 7.

6.4 Fake photons from misidentified jets

Background contributions from events in which a jet is misidentified as a photon are estimated using a sideband counting method [62]. This method relies on counting photon candidates in four regions of a two-dimensional space, defined by the isolation transverse energy and by the quality of the identification criteria. A signal region (region A) is defined by photon candidates that are isolated with tight identification. Three background regions are defined, consisting of photon candidates which are tight and non-isolated (region B), non-tight and isolated (region C) or non-tight and non-isolated (region D). The method relies on the assumption that the isolation profile in the non-tight region is the same as that of the background in the tight region. This has been verified in MC samples by checking that the correlation factor, calculated as $(N_A * N_D / N_B * N_C)$ is compatible with unity within uncertainties. The number of background candidates in the signal region (N_A) is calculated by taking the ratio of the two non-tight regions (N_C / N_D) multiplied by the number of candidates in the tight, non-isolated region (N_B). A correction to the method is added in order to take into account the leakage of real photon events to the three background regions. The systematic uncertainty of the method is evaluated by varying the criteria of tightness and isolation used to define the four regions. This estimate also accounts for the contribution from multi-jet events, which can mimic the $\gamma + E_T^{\text{miss}}$ signature if one jet is misreconstructed as a photon and one or more of the other jets are poorly reconstructed, resulting in large E_T^{miss} . This method is then used to evaluate the contribution of jets misidentified as photons in all analysis regions: the SRs and their associated four control regions, 1muCR, 2muCR, 2eleCR and PhJetCR. The estimated contribution from this background in the SRs and the associated uncertainty are reported in Sect. 7.

6.5 Final background estimation

The normalisation factors for the $W\gamma$, $Z\gamma$ and γ +jets backgrounds are obtained via a profile likelihood fit (referred to as the background-only fit). For this fit, a likelihood function is built as the product of Poisson probability functions of the observed and expected event yields in the control regions. The event yield in the corresponding SR is not considered. The systematic uncertainties, described in Sect. 8, are treated as Gaussian-distributed nuisance parameters in the likelihood function. For each CR, the inputs to the fit are: the number of events observed in the data, the expected numbers of $W/Z\gamma$

Table 3 Normalisation factors (scale factors k) obtained from a background-only inclusive-SR fit performed in each inclusive SR (the first three columns) and scale factors k' obtained from a background-only multiple-bin fit performed simultaneously in the three regions

| Signal region | E_T^{miss} [GeV] | $k_{W\gamma}$ | $k_{Z\gamma}$ | $k_{\gamma+\text{jets}}$ | $k'_{W\gamma}$ | $k'_{Z\gamma}$ | $k'_{\gamma+\text{jets}}$ |
|---------------|---------------------------|-----------------|-----------------|--------------------------|-----------------|-----------------|---------------------------|
| SRI1 | >150 | 1.05 ± 0.09 | 1.10 ± 0.09 | 1.07 ± 0.25 | | | |
| SRI2 | >225 | 1.04 ± 0.11 | 1.14 ± 0.13 | 1.06 ± 0.25 | | | |
| SRI3 | >300 | 1.04 ± 0.15 | 1.27 ± 0.23 | 1.06 ± 0.24 | 1.03 ± 0.14 | 1.27 ± 0.23 | |
| SRE1 | 150–225 | | | | 1.06 ± 0.10 | 1.10 ± 0.10 | 1.07 ± 0.25 |
| SRE2 | 225–300 | | | | 1.02 ± 0.12 | 1.09 ± 0.14 | |

SRE1, SRE2 and SRI3 (the last three columns), where $k'_{\gamma+\text{jets}}$ applies to all exclusive signal regions. The errors shown include both the statistical and systematic uncertainties

and γ +jets background events, which are taken from MC simulations and whose normalisations are free parameters in the fit, and the number of fake-photon events obtained from the data-driven techniques.

Two different configurations are used for the fit: the background-only inclusive fit, which determines the normalisations for $W\gamma$, $Z\gamma$ and γ +jets backgrounds for each inclusive SR independently and the background-only multiple-bin fit, which determines the normalisations for the three exclusive SRs simultaneously. In the former case, four CRs corresponding to a given SR are used to obtain the normalisations. In the latter case, all ten CRs (the three leptonic CRs for each SR and the PhJetCR) associated with the three exclusive SRs are used. These fits are described in more detail in the following.

6.5.1 Background-only inclusive-SR fit

Background estimates in each inclusive SR are derived from a simultaneous fit to the respective four control regions (1muCR, 2muCR, 2eleCR and PhJetCR). The fitted values of the normalisation factors for $W/Z\gamma$ and γ +jets backgrounds (scale factors k) are reported in Table 3. Although the PhJetCR is defined in the same way for all SRs, the $k_{\gamma+\text{jets}}$ factors in the three inclusive SRs differ slightly because they are fitted together with the other CRs, which are different for the different SRs.

The inclusive-SR fit is used to set model-independent limits, as shown in Sect. 9.

6.5.2 Background-only multiple-bin fit

A background-only multiple-bin fit is performed using simultaneously the control regions corresponding to the three signal regions SRE1, SRE2 and SRI3, which are mutually exclusive. The γ +jets normalisation factor is fixed in the common control region at low E_T^{miss} (PhJetCR), while the $W\gamma$ and $Z\gamma$ normalisation factors are fitted in each E_T^{miss} range separately. The estimated normalisation factors (scale factors k')

after this multiple-bin fit for each of the three SRs are also reported in Table 3. As expected, they agree within uncertainties with the scale factors k obtained from the inclusive-SR fit.

Post-fit distributions of E_T^{miss} in the four control regions are shown in Fig. 2. The scale factors k' from the multiple-bin fit are used for the different E_T^{miss} ranges to produce these figures. These distributions illustrate the contribution from the different background processes.

The multiple-bin fit is used to set exclusion limits in the models studied, if no excess is found in the data, as discussed in Sect. 9.

7 Results

Table 4 presents the observed number of events and the SM background predictions in SRI1 that is the most inclusive signal region with the lowest E_T^{miss} threshold, as obtained from the simultaneous inclusive-SR fit to its CRs. The corresponding number of events is shown in the three lepton CRs and in PhJetCR. For the SM predictions both the statistical and systematic uncertainties, described in Sect. 8, are included.

Table 5 shows the observed number of events and the total SM background prediction after the fit in all signal regions. For SRI1, SRI2 and SRI3 regions the expected SM event yields are obtained from the inclusive-SR fit to each region; for SRE1 and SRE2 regions the expected SM event yields are obtained from the multiple-bin fit to the regions SRE1, SRE2 and SRI3. The expected SM event yields in SRI3 are the same when obtained from the multiple-bin fit. The numbers of observed events in the corresponding lepton CRs for each SR are also reported.

The post-fit E_T^{miss} and E_T^γ distributions are shown in Fig. 3 after applying the scale factors k' from the multiple-bin fit. Only the E_T^{miss} distribution is used in the multiple-bin fit, as discussed in Sect. 9.

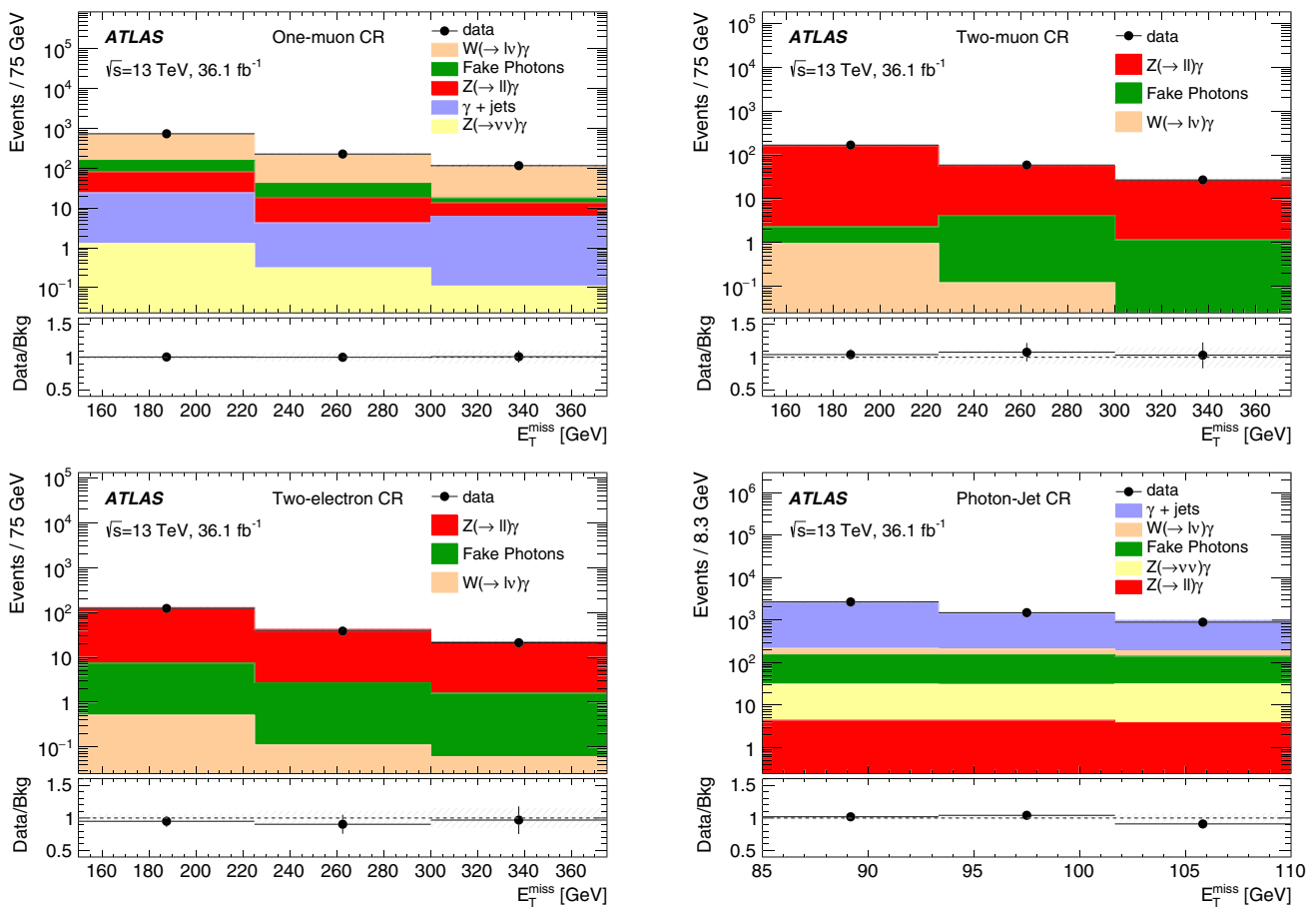


Fig. 2 Distribution of E_T^{miss} in data and for the expected total background in the CRs: 1muCR (top left), 2muCR (top right), 2eleCR (bottom left) and PhJetCR (bottom right). In 1muCR and 2muCR, the muons are treated as non-interacting particles in the E_T^{miss} reconstruction; the electrons are handled similarly in 2eleCR. The total background expectation is normalised using the scale factors k' derived from the multiple-

bin fit. For 1muCR, 2muCR and 2eleCR, overflows are included in the third bin. The error bars are statistical, and the dashed band includes statistical and systematic uncertainties determined by the multiple-bin fit. The lower panel shows the ratio of data to expected background event yields

Table 4 Observed event yields in 36.1 fb^{-1} of data compared to expected yields from SM backgrounds in the signal region SRI1 and in its four control regions (CRs), as predicted from the simultaneous fit to CRs of SRI1 (see text). The MC yields before the fit are also shown. The uncertainty includes both the statistical and systematic uncertainties described in Sect. 8. The uncertainty on the pre-fit background is the pre-fit uncertainty, while the uncertainties on the fitted background

are post-fit uncertainties. The latter are constrained by the fit as the use of control regions to normalise the dominant backgrounds allows to partially cancel some systematic uncertainties (see Sect. 8 for more details). The individual uncertainties can be correlated and do not necessarily add in quadrature to equal the total background uncertainty. The total fitted background does not match exactly the sum of the individual contributions because of the rounding applied

| | SRI1 | 1muCR | 2muCR | 2eleCR | PhJetCR |
|---------------------------------|----------------|---------------|-----------------|-----------------|-----------------|
| Observed events | 2400 | 1083 | 254 | 181 | 5064 |
| Fitted background | 2600 ± 160 | 1083 ± 33 | 243 ± 13 | 193 ± 10 | 5064 ± 80 |
| $Z(\rightarrow \nu\nu)\gamma$ | 1600 ± 110 | 1.7 ± 0.2 | – | – | 81 ± 6 |
| $W(\rightarrow \ell\nu)\gamma$ | 390 ± 24 | 866 ± 40 | 1.1 ± 0.3 | 0.7 ± 0.1 | 163 ± 9 |
| $Z(\rightarrow \ell\ell)\gamma$ | 35 ± 3 | 77 ± 5 | 233 ± 13 | 180 ± 10 | 13 ± 1 |
| γ + jets | 248 ± 80 | 33 ± 8 | – | – | 4451 ± 80 |
| Fake photons from electrons | 199 ± 40 | 17 ± 3 | 0.50 ± 0.13 | 0.09 ± 0.04 | 72 ± 14 |
| Fake photons from jets | 152 ± 22 | 88 ± 19 | 7.9 ± 3.8 | 12 ± 5 | 284 ± 28 |
| Pre-fit background | 2400 ± 200 | 1025 ± 72 | 218 ± 15 | 181 ± 13 | 4800 ± 1000 |

Table 5 Observed event yields in 36.1 fb^{-1} of data compared to expected yields from SM backgrounds in all signal regions, as predicted from the simultaneous fit to their respective CRs (see text). The first three columns report the yields obtained from the inclusive-SR fit, while the two last columns report the yields obtained from the multiple-bin fit. The uncertainty includes both the statistical and systematic uncertainties described in Sect. 8. The uncertainties are post-fit uncertainties and are constrained by the fit as the use of control regions to normalise

| | SRI1 | SRI2 | SRI3 | SRE1 | SRE2 |
|---------------------------------|----------------|---------------|--------------------|----------------|---------------|
| Observed events | 2400 | 729 | 236 | 1671 | 493 |
| Fitted background | 2600 ± 160 | 765 ± 59 | 273 ± 37 | 1900 ± 140 | 501 ± 44 |
| $Z(\rightarrow \nu\nu)\gamma$ | 1600 ± 110 | 543 ± 54 | 210 ± 35 | 1078 ± 89 | 342 ± 41 |
| $W(\rightarrow \ell\nu)\gamma$ | 390 ± 24 | 109 ± 9 | 33 ± 4 | 282 ± 22 | 75 ± 8 |
| $Z(\rightarrow \ell\ell)\gamma$ | 35 ± 3 | 7.8 ± 0.8 | 2.2 ± 0.4 | 27 ± 3 | 5.7 ± 0.7 |
| γ + jets | 248 ± 80 | 22 ± 7 | 5.2 ± 1.0 | 225 ± 80 | 17 ± 6 |
| Fake photons from electrons | 199 ± 40 | 47 ± 11 | 13 ± 3 | 152 ± 28 | 34 ± 8 |
| Fake photons from jets | 152 ± 22 | 37 ± 15 | $9.7_{-9.7}^{+10}$ | 115 ± 24 | 27 ± 9 |
| Observed events in 1muCR | 1083 | 343 | 116 | 740 | 227 |
| Observed events in 2muCR | 254 | 86 | 27 | 168 | 59 |
| Observed events in 2eleCR | 181 | 59 | 21 | 122 | 38 |
| Observed events in PhJetCR | 5064 | 5064 | 5064 | 5064 | 5064 |

the dominant backgrounds allows to partially cancel some systematic uncertainties (see Sect. 8 for more details). The individual uncertainties can be correlated and do not necessarily add in quadrature to equal the total background uncertainty. The observed number of events in the four CRs relative to each SR is also shown. The total fitted background does not match exactly the sum of the individual contributions because of the rounding applied

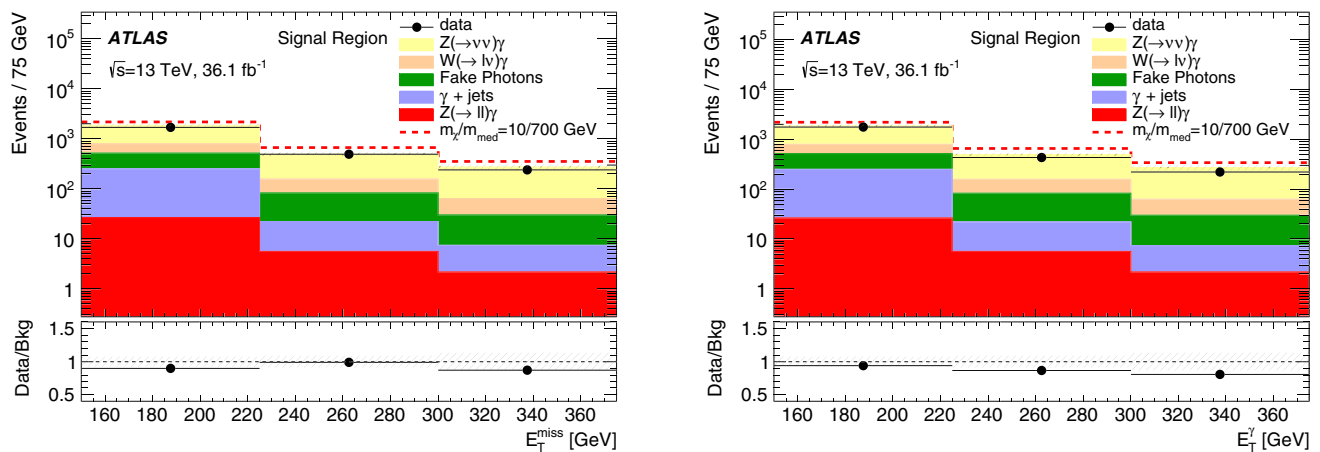


Fig. 3 Distribution of E_T^{miss} (left) and of E_T^γ (right) in the signal regions for data and for the expected total background; the total background expectation is normalised using the scale factors k' derived from the multiple-bin fit. Overflows are included in the third bin. The error bars are statistical, and the dashed band includes statistical and systematic

uncertainties determined by the fit. The expected yield of events from the simplified model with $m_\chi = 10 \text{ GeV}$ and an axial-vector mediator of mass $m_{\text{med}} = 700 \text{ GeV}$ with $g_q = 0.25$ and $g_\chi = 1.0$ is stacked on top of the background prediction. The lower panel shows the ratio of data to expected background event yields

8 Systematic uncertainties

The systematic uncertainties are treated as Gaussian-distributed nuisance parameters in the likelihood function fitted to obtain the final background predictions in the SRs, as described in Sect. 6. The fit provides constraints on many sources of systematic uncertainty, as the normalisations of the dominant background processes are fitted parameters; only the

uncertainties affecting the extrapolation between CRs and SRs therefore remain important.

The fitted uncertainties are presented as percentages of the total background predictions in SRs. The total background prediction uncertainty, including systematic and statistical contributions, varies from 6.1 to 14% for the five SRs, dominated by the statistical uncertainty in the control regions, which varies from approximately 4.3 to 10%.

Table 6 Breakdown of the relevant uncertainties in the background estimates for all SRs. The uncertainties are given relative to the expected total background yield after the inclusive-SR fit for SRI1, SRI2 and SRI3 and after the multiple-bin fit for SRE1 and SRE2. The total statis-

tical uncertainty is dominated by the statistical uncertainty in CRs. The individual uncertainties can be correlated and do not necessarily add in quadrature to equal the total background uncertainty

| | SRI1 | SRI2 | SRI3 | SRE1 | SRE2 |
|---|-------|------|-------|-------|-------|
| Total background | 2600 | 765 | 273 | 1900 | 501 |
| Total (statistical+systematic) uncertainty | 6.1% | 7.7% | 14% | 7.7% | 8.8% |
| Statistical uncertainty only | 4.3% | 6.2% | 10% | 5.5% | 7.8% |
| Jet fake rate (Sect. 6.4) | 1.3% | 3.0% | 5.3% | 1.7% | 3.3% |
| Electron fake rate (Sect. 6.3) | 1.5% | 1.5% | 1.2% | 1.5% | 1.6% |
| Jet energy scale [56] | 4.1% | 1.9% | 1.4% | 5.6% | 0.6% |
| Jet energy resolution [69] | 0.7% | 0.2% | 0.5% | 0.8% | 0.3% |
| E_T^{miss} soft-term scale and resolution [61] | 0.9% | 0.4% | 0.7% | 1.1% | 1.0% |
| Muon reconstruction/isolation efficiency [53] | 1.4% | 1.3% | 1.6% | 1.3% | 1.4% |
| Electron reco/identification/isolation efficiency [70] | 1.0% | 1.3% | 1.3% | 0.8% | 1.2% |
| Electron and photon energy scale [51] | 0.2% | 0.5% | 0.4% | <0.1% | <0.1% |
| Electron and photon energy resolution [51] | <0.1% | 0.3% | 0.2% | 0.1% | 1.0% |
| Photon efficiency [52] | 0.1% | 1.0% | <0.1% | 0.2% | <0.1% |
| γ +jets modelling | 1.5% | 0.3% | 0.3% | 2.3% | 0.4% |
| $\langle\mu\rangle$ distribution in MC simulation (Sect. 3) | 1.3% | 0.3% | 0.9% | 1.7% | 0.3% |

The relevant uncertainties (giving a contribution of more than 0.1% in at least one SR) are summarised in Table 6 for all SRs.

Aside from the uncertainty due to the statistical precision from the CRs, the largest relative systematic uncertainties are due to the uncertainty in the rate of fake photons from jets, which contributes 1.3–5.3%, increasing for SRI3 and SRE2 because of the smaller number of events available for the estimation, and to the uncertainty in the jet energy scale, which contributes 1.4–5.6%, decreasing in the regions with larger E_T^{miss} . The systematic uncertainty due to the modelling of the γ +jets background, which affects the extrapolation of this background from the PhJetCR to the SRs, is evaluated by independently varying the following four parameters with respect to the nominal values used in the MC generation: the renormalisation, factorisation and resummation scales by factors of 2.0 and 0.5, and the CKKW matching scale [65] to 15 and 30 GeV (the nominal value being 20 GeV). For the $W/Z\gamma$ backgrounds, the lepton identification/reconstruction efficiency uncertainties are propagated from the leptonic CRs to the SRs in terms of veto efficiency uncertainties. After the fit, the uncertainty in the luminosity [66] is found to have a negligible impact on the background estimation. The uncertainties due to the PDF have an impact on the $W/Z\gamma$ samples in each region but the effect on normalisation is largely absorbed in the fit, so their impact is negligible.

For the signal-related systematic uncertainties, the uncertainties due to PDF are evaluated following the PDF4LHC recommendations [67] and using a reweighting procedure implemented in the LHAPDF Tool [68], while uncertainties

due to the scales are evaluated by varying the renormalisation and factorisation scales by factors of 2.0 and 0.5 with respect to the nominal values used in the MC generation. The uncertainties in initial- and final-state radiation, due to the choice of parton shower and multiple-parton-interaction parameters used with PYTHIA 8.186 are estimated by generating MC samples with the alternative tunes described in Ref. [31]. The PDF, scale and tune each induce uncertainties of up to about 5% in the acceptance (and cross section) in the simplified DM models.

9 Interpretation of results

The event yields observed in data are consistent within uncertainties with the predicted SM background event yields in all inclusive SRs, as shown in Table 5. The results from the SRs shown in Sect. 7 are therefore interpreted in terms of exclusion limits in models that would produce an excess of $\gamma + E_T^{\text{miss}}$ events. Upper bounds are calculated using a one-sided profile likelihood ratio and the CL_S technique [71, 72], evaluated using the asymptotic approximation [73]. The likelihood fit includes both the SRs and their CRs.

The upper limits on the visible cross section, defined as the product of the cross section times the acceptance times the reconstruction efficiency defined in a fiducial region, $\sigma \times A \times \epsilon$, of a potential BSM signal, are obtained from the three inclusive SRs. The value of A for a particular model is computed by applying the same selection criteria as in the SR but at the particle level; in this computation E_T^{miss}

Table 7 The observed and expected limits at 95% CL on the fiducial cross section $\sigma \times A$. The values of the fiducial reconstruction efficiency (ϵ), which is used for the calculation of the fiducial cross section, and of the acceptance (A) are also shown

| Region | $\sigma \times A$ limit [fb] | | |
|-----------------------------------|------------------------------|---------------|---------------|
| | SRI1 | SRI2 | SRI3 |
| 95% CL observed | 7.0 | 3.7 | 2.3 |
| 95% CL expected | 10.6 | 4.5 | 3.0 |
| 95% CL expected ($\pm 1\sigma$) | 14.5 ± 7.7 | 6.2 ± 3.3 | 4.2 ± 2.2 |
| A [%] | 14–48 | 5–31 | 2–19 |
| ϵ [%] | 84–95 | 73–86 | 64–85 |

is given by the vector sum of the transverse momenta of all non-interacting particles. The A values with the selection for SRI1 or SRI2 or SRI3 are reported in Table 7 for the simplified DM models; the lowest values are found in models with low-mass off-shell mediators and the highest values in models with high-mass on-shell mediators. The observed and expected upper limits, at 95% confidence level (CL), on the fiducial cross section, defined as $\sigma \times A$ are shown in Table 7. They are computed by dividing the limit on the visible cross section by the fiducial reconstruction efficiency ϵ shown in the same table; as in the case of the acceptance, the efficiency is smaller in the high- E_T^{miss} bins. The lowest efficiency for each signal region is used in a conservative way to set the fiducial cross-section limit. These limits can be extrapolated to models producing $\gamma + E_T^{\text{miss}}$ events once A is known, assuming that the conservative efficiency applies.

The expected limit on the signal strength in the simplified DM model is computed with the inclusive-SR fit for the various inclusive regions and with the multiple-bin fit in order to determine which strategy to adopt for limit setting. While SRI1 is the inclusive SR that gives the most stringent expected limits at very low DM/mediator masses, SRI2 is the inclusive SR providing the most stringent limits in the rest of the parameter space; SRI3, which has larger uncertainties, is not able to set better expected constraints on high-mass models in spite of their harder E_T^{miss} spectra. The multiple-bin fit, making use of the expected signal distribution in E_T^{miss} by combining the information from the various exclusive SRs,

allows more stringent expected limits to be set than in any of the inclusive signal regions.

The results are presented for both the axial-vector and vector mediators using different couplings to show the complementarity to the direct searches in $X + E_T^{\text{miss}}$ events and the searches looking for the mediator, such as dijet or dilepton resonance searches, as recommended in Ref. [74]. Four models are considered with different mediators and different couplings to quarks, to DM particles, and to leptons, and these models are summarised in Table 8. As the choices of mediators and of couplings only affect the signal cross section and not the acceptance for signal events, the events generated for the axial-vector mediator with $g_q = 0.25$, $g_\chi = 1$ and $g_\ell = 0$ (model A1), described in Sect. 3, can be re-scaled in order to obtain results for the other three models.

When placing limits in specific models, the signal-related systematic uncertainties estimated as described in Sect. 8 affecting $A \times \epsilon$ (PDF, scales, initial- and final-state radiation) are included in the statistical analysis, while the uncertainties affecting the cross section (PDF, scales) are indicated as bands around the observed limits and written as σ_{theo} .

Simplified models with explicit mediators are valid for all values of momentum transfer in pp collisions [10]. Figure 4 (top left) shows the observed and expected contours corresponding to a 95% CL exclusion as a function of m_{med} and m_χ for the simplified model A1. The region of the plane under the limit curves is excluded. The region not allowed due to perturbative unitarity violation is to the left of the line defined by $m_\chi = \sqrt{\pi/2}m_{\text{med}}$ [75]. The line corresponding to the DM thermal relic abundance measured by the Planck collaboration [76] is also indicated in the figure; it is obtained as detailed in Ref. [74]. Figure 4 (top right) shows the contours for the A2 model, while Fig. 4 (bottom left) and (bottom right) show the contours for the V1 and V2 models, respectively. The search excludes mediator masses below the values reported in Table 8 for χ masses below the values reported in the same table. The limits for the model A1 are more stringent than the limits obtained with the 2015 data only [17], which excluded mediator masses below 710 GeV for χ masses below 150 GeV.

Figure 5 (left) shows the contours corresponding to a 90% CL exclusion translated into the plane of χ -proton spin-dependent (SD) scattering cross sections vs. m_χ for the

Table 8 Observed limits at 95% CL on the mediator mass and the DM particle mass for the four models considered. The mediators and couplings to quarks, to dark-matter particles and to leptons are specified for each model

| Model | Mediator | g_q | g_χ | g_ℓ | Limit on m_{med} [GeV] for low m_χ | Limit on m_χ [GeV] reaching as high as |
|-------|--------------|-------|----------|----------|--|---|
| A1 | Axial-vector | 0.25 | 1 | 0 | 1200 | 340 |
| A2 | Axial-vector | 0.1 | 1 | 0.1 | 750 | 230 |
| V1 | Vector | 0.25 | 1 | 0 | 1200 | 480 |
| V2 | Vector | 0.1 | 1 | 0.01 | 750 | 320 |

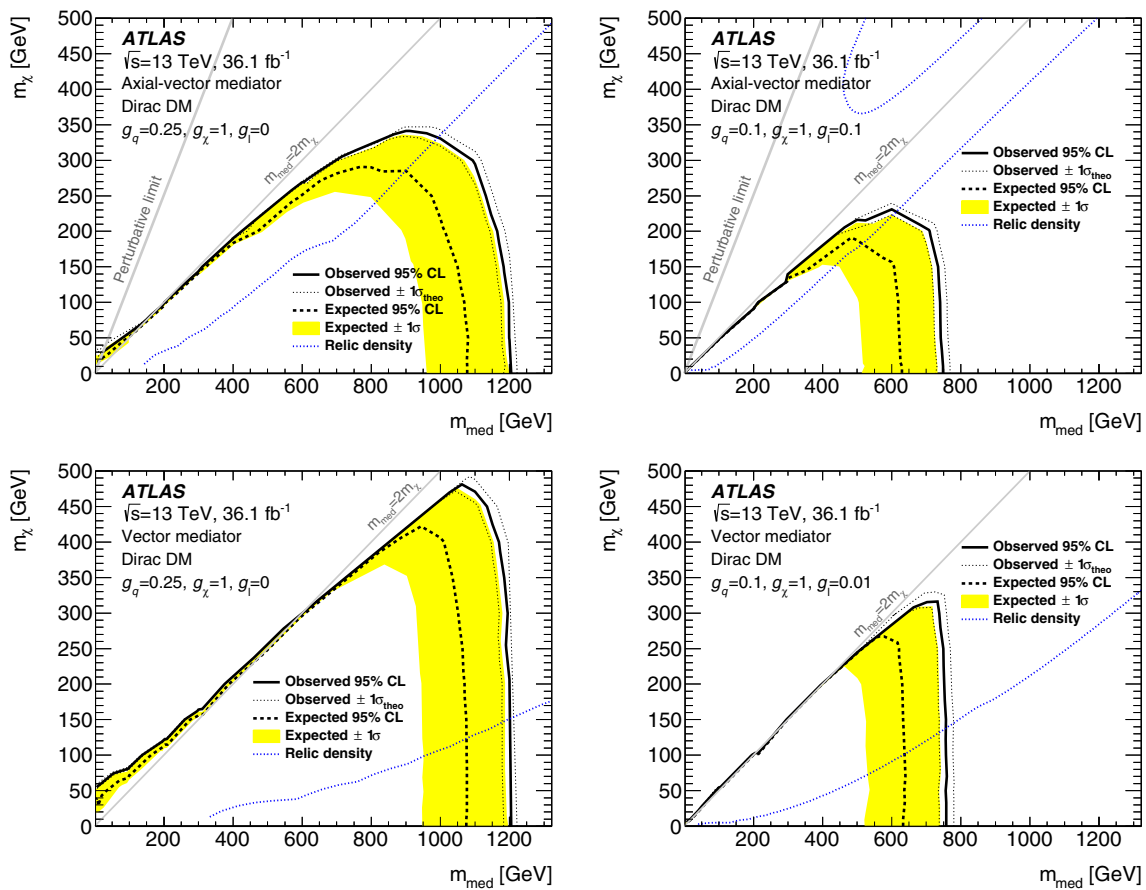


Fig. 4 The observed and expected 95% CL exclusion contours for a simplified model of dark-matter production involving an axial-vector operator, Dirac DM and couplings $g_q = 0.25$, $g_\chi = 1$ and $g_\ell = 0$ as a function of the dark-matter mass m_χ and the mediator mass m_{med} (upper left). The plane under the limit curve is excluded. The same is shown for an axial-vector operator with couplings $g_q = 0.1$, $g_\chi = 1$ and $g_\ell = 0.1$ (top right), for a vector operator with couplings $g_q = 0.25$, $g_\chi = 1$ and $g_\ell = 0$ (bottom left) and for a vector operator with couplings

$g_q = 0.1$, $g_\chi = 1$ and $g_\ell = 0.01$ (bottom right). The region on the left is excluded by the perturbative limit which is relevant for axial-vector mediators [77]. The relic density curve [74, 76] is also shown: at higher mediator masses, the DM would be overabundant; at lower values, it would be underabundant; for the axial-vector scenario shown in the upper right figure, the region above the relic density curve at high dark-matter masses is also overabundant

axial-vector mediator model A1. Bounds on the χ -proton cross section are obtained following the procedure described in Ref. [77], assuming that the axial-vector mediator with couplings as in A1 is solely responsible for both collider χ pair production and for χ -proton scattering. In this plane, a comparison with the result from direct DM searches [78, 79] is possible. The limit placed in this search extends to arbitrarily low values of m_χ , as the acceptance at lower mass values is the same as the one at the lowest m_χ value shown here. The search provides stringent limits on the scattering cross section of the order of 10^{-42} cm^2 up to m_χ masses of about 300 GeV. These results allow complementary limits to be set on the χ -proton scattering cross section in the low DM mass region where the direct DM search experiments have less sensitivity due to the very low energy recoils that such low-mass dark-matter particles would induce. Figure 5

(right) shows the limit contours in the plane of the χ -nucleon spin-independent (SI) scattering cross section vs. m_χ for the vector mediator model V1 compared with results of direct DM searches [80–83]. In this case the limit on the scattering cross section is of the order of 10^{-41} cm^2 up to m_χ masses of about 500 GeV.

In the case of the model of $\gamma\gamma\chi\bar{\chi}$ interactions, lower limits are placed on the effective mass scale M_* as a function of m_χ , as shown in Fig. 6. In this model, which presents a hard E_T^{miss} spectrum, the signal events mainly contribute to the $E_T^{\text{miss}} > 300 \text{ GeV}$ bin. The search excludes model values of M_* up to about 790 GeV, which is a more stringent limit than the one placed in earlier searches [17]. The EFT description is not always valid at these scales. The effect of the truncation for two representative values of the EFT coupling, g^* , is shown in the same figure, assuming that the scale at which

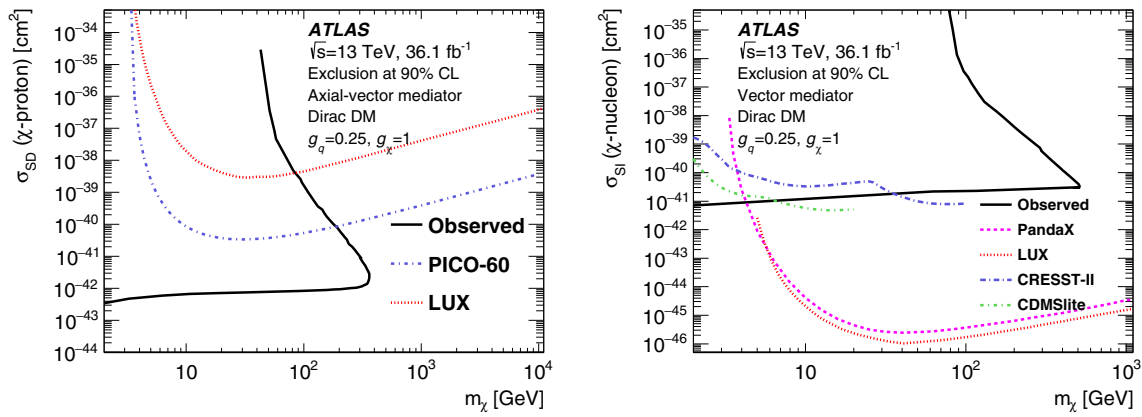


Fig. 5 The 90% CL exclusion limit on the χ -proton scattering cross section in a simplified model of dark-matter production involving an axial-vector operator, Dirac DM and couplings $g_q = 0.25, g_\chi = 1$ and $g_\ell = 0$ as a function of the dark-matter mass m_χ . Also shown are results at 90% CL from two direct dark-matter search experiments [78, 79] (left). The 90% CL exclusion limit on the χ -nucleon scattering

cross section in a simplified model of dark-matter production involving a vector operator, Dirac DM and couplings $g_q = 0.25, g_\chi = 1$ and $g_\ell = 0$ as a function of the dark-matter mass m_χ (right); also shown are results at 90% CL from four direct dark-matter search experiments [80–83]

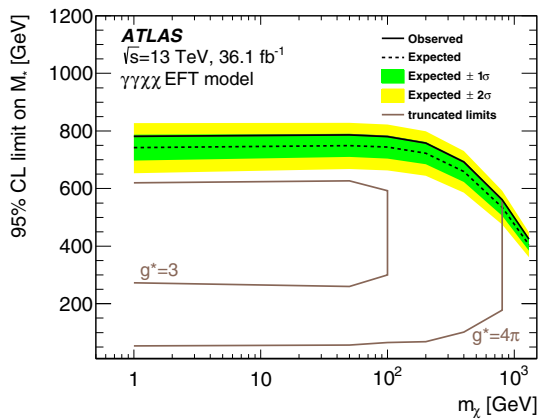


Fig. 6 The observed and expected 95% CL limits on M_* for a dimension-7 operator EFT model with a contact interaction of type $\gamma\gamma\chi\chi$ as a function of dark-matter mass m_χ . Results where EFT truncation is applied are also shown, assuming representative coupling values, g^* , of 3 and 4π : for the maximal coupling value of 4π , the truncation has almost no effect; for lower coupling values, the exclusion limits are confined to a smaller area of the parameter space

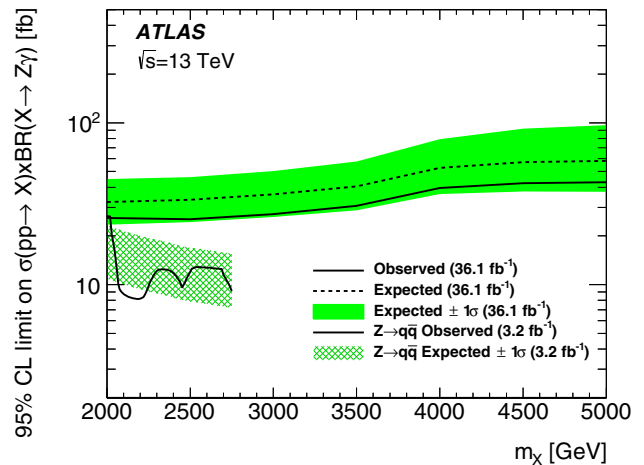


Fig. 7 The observed (expected) limit at 95% CL on the production cross section of a $Z\gamma$ resonance as a function of its mass. The limits from the search in the $Z \rightarrow q\bar{q}$ channel with 3.2 fb^{-1} [84] are also reported

the EFT description becomes invalid (M_{cut}) is related to M_* through $M_{\text{cut}} = g^* M_*$. For the maximal coupling value of 4π , the truncation has almost no effect; for lower coupling values, the exclusion limits are confined to a smaller area of the parameter space.

The results are also interpreted in terms of a limit on the cross section for the production of a narrow heavy scalar $Z\gamma$ resonance produced through gluon–gluon fusion. Figure 7 shows the observed and expected limit at 95% CL on the production cross section of a $Z\gamma$ resonance as a function of its mass. The limit is produced in exactly the same way as the other signal samples, where an excess of events is sought in

the three exclusive signal regions by using the multiple-bin fit. The heavy resonances are expected to populate mainly the $E_T^{\text{miss}} > 300 \text{ GeV}$ signal region as they would have a hard E_T^{miss} spectrum. The upper bound on $m_{\ell\ell\gamma}$ applied in 2eleCR and 2muCR (see Sect. 6.1) suppresses the contamination from potential high-mass $Z\gamma$ resonances in these control regions. Limits on such a resonance were also placed by bump searches in the very sensitive dileptonic channel and the hadronic channel for masses below and above 1.5 TeV, respectively [84]. Although the Z boson branching ratio to neutrinos is higher than to charged leptons, the presence of E_T^{miss} makes the search in this channel much less sensitive than in the dileptonic channel; the region of interest for the

analysis discussed here lies at higher masses, where it can complement the searches using Z boson hadronic decays whose limits, obtained with 3.2 fb^{-1} , are reported in the same figure. The observed (expected) limits at 95% CL on the production of a $Z\gamma$ resonance are 26 and 43 fb (32 and 58 fb) for masses of 2 and 5 TeV, respectively.

10 Conclusion

Results are reported from a search for dark matter in events with a high transverse energy photon and large missing transverse momentum in pp collisions at $\sqrt{s} = 13 \text{ TeV}$ at the LHC. Data collected by the ATLAS experiment and corresponding to an integrated luminosity of 36.1 fb^{-1} are used. The observed data are consistent with the Standard Model expectations. The observed (expected) upper limits on the fiducial cross section for the production of events with a photon and large missing transverse momentum are 7.0 and 2.3 fb (10.6 and 3.0 fb) at 95% CL for E_T^{miss} thresholds of 150 and 300 GeV, respectively. For the simplified dark-matter model considered, the search excludes axial-vector and vector mediators with masses below 750–1200 GeV for χ masses below 230–480 GeV at 95% CL, depending on the couplings chosen. For an EFT $\gamma\gamma\chi\bar{\chi}$ model of dark-matter production, values of the suppression scale M_* up to 790 GeV are excluded at 95% CL and the effect of truncation for various coupling values is reported. The observed (expected) limits at 95% CL on the production cross section for a narrow $Z\gamma$ scalar resonance are 26 and 43 fb (32 and 58 fb) for masses of 2 and 5 TeV, respectively.

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References

- OPAL Collaboration, G. Abbiendi et al., Photonic events with missing energy in e^+e^- collisions at $\sqrt{s} = 189 \text{ GeV}$. *Eur. Phys. J. C* **18**, 253 (2000). doi:[10.1007/s100520000522](https://doi.org/10.1007/s100520000522). arXiv:[hep-ex/0005002](https://arxiv.org/abs/hep-ex/0005002)
- L3 Collaboration, P. Achard et al., Single photon and multi-photon events with missing energy in e^+e^- collisions at LEP. *Phys. Lett. B* **587**, 16 (2004). doi:[10.1016/j.physletb.2004.01.010](https://doi.org/10.1016/j.physletb.2004.01.010). arXiv:[hep-ex/0402002](https://arxiv.org/abs/hep-ex/0402002)
- DELPHI Collaboration, J. Abdallah et al., Photon events with missing energy in e^+e^- collisions at $\sqrt{s} = 130 \text{ GeV}$ to 209 GeV . *Eur. Phys. J. C* **38**, 395 (2005). doi:[10.1140/epjc/s2004-02051-8](https://doi.org/10.1140/epjc/s2004-02051-8). arXiv:[hep-ex/0406019](https://arxiv.org/abs/hep-ex/0406019)
- CDF Collaboration, T. Aaltonen et al., Search for large extra dimensions in final states containing one photon or jet and large missing transverse energy produced in $p\bar{p}$ collisions at $\sqrt{s} = 1.96 \text{ TeV}$. *Phys. Rev. Lett.* **101**, 181602 (2008). doi:[10.1103/PhysRevLett.101.181602](https://doi.org/10.1103/PhysRevLett.101.181602). arXiv:[0807.3132](https://arxiv.org/abs/0807.3132) [hep-ex]
- D0 Collaboration, V.M. Abazov et al., Search for large extra dimensions via single photon plus missing energy final states at $\sqrt{s} = 1.96 \text{ TeV}$. *Phys. Rev. Lett.* **101**, 011601 (2008). doi:[10.1103/PhysRevLett.101.011601](https://doi.org/10.1103/PhysRevLett.101.011601). arXiv:[0803.2137](https://arxiv.org/abs/0803.2137) [hep-ex]
- G. Bertone, D. Hooper, J. Silk, Particle dark matter: evidence, candidates and constraints. *Phys. Rep.* **405**, 279 (2005). doi:[10.1016/j.physrep.2004.08.031](https://doi.org/10.1016/j.physrep.2004.08.031). arXiv:[hep-ph/0404175](https://arxiv.org/abs/hep-ph/0404175)
- J. Goodman et al., Constraints on dark matter from colliders. *Phys. Rev. D* **82**, 116010 (2010). doi:[10.1103/PhysRevD.82.116010](https://doi.org/10.1103/PhysRevD.82.116010). arXiv:[1008.1783](https://arxiv.org/abs/1008.1783) [hep-ph]
- J. Abdallah et al., Simplified models for dark matter searches at the LHC. *Phys. Dark Univ.* **9–10**, 8 (2015). doi:[10.1016/j.dark.2015.08.001](https://doi.org/10.1016/j.dark.2015.08.001). arXiv:[1506.03116](https://arxiv.org/abs/1506.03116) [hep-ph]
- O. Buchmueller, M.J. Dolan, S.A. Malik, C. McCabe, Characterising dark matter searches at colliders and direct detection experiments: vector mediators. *JHEP* **01**, 037 (2015). doi:[10.1007/JHEP01\(2015\)037](https://doi.org/10.1007/JHEP01(2015)037). arXiv:[1407.8257](https://arxiv.org/abs/1407.8257) [hep-ph]
- D. Abercrombie et al., Dark matter benchmark models for early LHC run-2 searches: report of the ATLAS/CMS dark matter forum (2015). arXiv:[1507.00966](https://arxiv.org/abs/1507.00966) [hep-ex]
- G. Busoni, A. De Simone, E. Morgante, A. Riotto, On the validity of the effective field theory for dark matter searches at the LHC. *Phys. Lett. B* **728**, 412 (2014). doi:[10.1016/j.physletb.2013.11.069](https://doi.org/10.1016/j.physletb.2013.11.069). arXiv:[1307.2253](https://arxiv.org/abs/1307.2253) [hep-ph]
- A. Crivellin, U. Haisch, A. Hibbs, LHC constraints on gauge boson couplings to dark matter. *Phys. Rev. D* **91**, 074028 (2015). doi:[10.1103/PhysRevD.91.074028](https://doi.org/10.1103/PhysRevD.91.074028). arXiv:[1501.00907](https://arxiv.org/abs/1501.00907) [hep-hp]

13. E. Eichten, K. Lane, Low-scale technicolor at the Tevatron and LHC. *Phys. Lett. B* **669**, 235 (2008). doi:[10.1016/j.physletb.2008.09.047](https://doi.org/10.1016/j.physletb.2008.09.047). arXiv:[0706.2339](https://arxiv.org/abs/0706.2339) [hep-ph]
14. I. Low, J. Lykken, G. Shaughnessy, Singlet scalars as Higgs imposters at the large hadron collider. *Phys. Rev. D* **84**, 035027 (2011). doi:[10.1103/PhysRevD.84.035027](https://doi.org/10.1103/PhysRevD.84.035027). arXiv:[1105.4587](https://arxiv.org/abs/1105.4587) [hep-ph]
15. ATLAS Collaboration, Search for dark matter candidates and large extra dimensions in events with a photon and missing transverse momentum in pp collision data at $\sqrt{s} = 7$ TeV with the ATLAS detector. *Phys. Rev. Lett.* **110**, 011802 (2013). doi:[10.1103/PhysRevLett.110.011802](https://doi.org/10.1103/PhysRevLett.110.011802). arXiv:[1209.4625](https://arxiv.org/abs/1209.4625) [hep-ex]
16. ATLAS Collaboration, Search for new phenomena in events with a photon and missing transverse momentum in pp collisions at $\sqrt{s} = 8$ TeV with the ATLAS detector. *Phys. Rev. D* **91**, 012008 (2015). doi:[10.1103/PhysRevD.91.012008](https://doi.org/10.1103/PhysRevD.91.012008). arXiv:[1411.1559](https://arxiv.org/abs/1411.1559) [hep-ex] [Erratum: *Phys. Rev. D* **92**, 059903 (2015)]
17. ATLAS Collaboration, Search for new phenomena in events with a photon and missing transverse momentum in pp collisions at $\sqrt{s} = 13$ TeV with the ATLAS detector. *JHEP* **06**, 059 (2016). doi:[10.1007/JHEP06\(2016\)059](https://doi.org/10.1007/JHEP06(2016)059). arXiv:[1604.01306](https://arxiv.org/abs/1604.01306) [hep-ex]
18. CMS Collaboration, Search for dark matter and large extra dimensions in pp collisions yielding a photon and missing transverse energy. *Phys. Rev. Lett.* **108**, 261803 (2012). doi:[10.1103/PhysRevLett.108.261803](https://doi.org/10.1103/PhysRevLett.108.261803). arXiv:[1204.0821](https://arxiv.org/abs/1204.0821) [hep-ex]
19. CMS Collaboration, Search for new phenomena in monophoton final states in proton–proton collisions at $\sqrt{s} = 8$ TeV. *Phys. Lett. B* **755**, 102–124 (2016). doi:[10.1016/j.physletb.2016.01.057](https://doi.org/10.1016/j.physletb.2016.01.057). arXiv:[1410.8812](https://arxiv.org/abs/1410.8812) [hep-ex]
20. ATLAS Collaboration, Search for new phenomena in final states with an energetic jet and large missing transverse momentum in pp collisions at $\sqrt{s} = 13$ TeV using the ATLAS detector. *Phys. Rev. D* **94**, 032005 (2016). doi:[10.1103/PhysRevD.94.032005](https://doi.org/10.1103/PhysRevD.94.032005). arXiv:[1604.07773](https://arxiv.org/abs/1604.07773) [hep-ex]
21. ATLAS Collaboration, Search for dark matter produced in association with a hadronically decaying vector boson in pp collisions at $\sqrt{s} = 13$ TeV with the ATLAS detector. *Phys. Lett. B* **763**, 251 (2016). doi:[10.1016/j.physletb.2016.10.042](https://doi.org/10.1016/j.physletb.2016.10.042). arXiv:[1608.02372](https://arxiv.org/abs/1608.02372) [hep-ex]
22. ATLAS Collaboration, Search for dark matter in association with a Higgs boson decaying to b -quarks in pp collisions at $\sqrt{s} = 13$ TeV with the ATLAS detector. *Phys. Lett. B* **765**, 11 (2017). doi:[10.1016/j.physletb.2016.11.035](https://doi.org/10.1016/j.physletb.2016.11.035). arXiv:[1609.04572](https://arxiv.org/abs/1609.04572) [hep-ex]
23. CMS Collaboration, Search for dark matter produced with an energetic jet or a hadronically decaying W or Z boson at $\sqrt{s} = 13$ TeV (2017). arXiv:[1703.01651](https://arxiv.org/abs/1703.01651) [hep-ex]
24. CMS Collaboration, Search for associated production of dark matter with a Higgs boson decaying to $b\bar{b}$ or $\gamma\gamma$ at $\sqrt{s} = 13$ TeV (2017). arXiv:[1703.05236](https://arxiv.org/abs/1703.05236) [hep-ex]
25. ATLAS Collaboration, The ATLAS experiment at the CERN large hadron collider. *JINST* **3**, S08003 (2008). doi:[10.1088/1748-0221/3/08/S08003](https://doi.org/10.1088/1748-0221/3/08/S08003)
26. ATLAS Collaboration, ATLAS insertable B-layer technical design report. ATLAS-TDR-19 (2010). <https://cds.cern.ch/record/1291633>
27. ATLAS Collaboration, ATLAS insertable B-layer technical design report addendum. ATLAS-TDR-19-ADD-1 (2012). <https://cds.cern.ch/record/1451888>
28. ATLAS Collaboration, Performance of the ATLAS trigger system in 2015 (2016). arXiv:[1611.09661](https://arxiv.org/abs/1611.09661) [hep-ex]
29. J. Alwall et al., The automated computation of tree-level and next-to-leading order differential cross sections, and their matching to parton shower simulations. *JHEP* **07**, 079 (2014). doi:[10.1007/JHEP07\(2014\)079](https://doi.org/10.1007/JHEP07(2014)079). arXiv:[1405.0301](https://arxiv.org/abs/1405.0301) [hep-ph]
30. T. Sjöstrand, S. Mrenna, P.Z. Skands, A brief introduction to PYTHIA 8.1. *Comput. Phys. Commun.* **178**, 852 (2008). doi:[10.1016/j.cpc.2008.01.036](https://doi.org/10.1016/j.cpc.2008.01.036). arXiv:[0710.3820](https://arxiv.org/abs/0710.3820) [hep-ph]
31. ATLAS Collaboration, ATLAS Pythia 8 tunes to 7 TeV data. ATLAS-PHYS-PUB-2014-021 (2014). <https://cds.cern.ch/record/1966419>
32. NNPDF Collaboration, R.D. Ball et al., Parton distributions for the LHC run II. *JHEP* **04**, 040 (2015). doi:[10.1007/JHEP04\(2015\)040](https://doi.org/10.1007/JHEP04(2015)040). arXiv:[1410.8849](https://arxiv.org/abs/1410.8849) [hep-ph]
33. M. Backovic et al., Higher-order QCD predictions for dark matter production at the LHC in simplified models with s-channel mediators. *Eur. Phys. J. C* **75**, 482 (2015). doi:[10.1140/epjc/s10052-015-3700-6](https://doi.org/10.1140/epjc/s10052-015-3700-6). arXiv:[1508.05327](https://arxiv.org/abs/1508.05327) [hep-ph]
34. S. Alioli, P. Nason, C. Oleari, E. Re, A general framework for implementing NLO calculations in shower Monte Carlo programs: the POWHEG BOX. *JHEP* **06**, 043 (2010). doi:[10.1007/JHEP06\(2010\)043](https://doi.org/10.1007/JHEP06(2010)043). arXiv:[1002.2581](https://arxiv.org/abs/1002.2581) [hep-ph]
35. H.-L. Lai et al., New parton distributions for collider physics. *Phys. Rev. D* **82**, 074024 (2010). doi:[10.1103/PhysRevD.82.074024](https://doi.org/10.1103/PhysRevD.82.074024). arXiv:[1007.2241](https://arxiv.org/abs/1007.2241) [hep-ph]
36. ATLAS Collaboration, Measurement of the Z/γ^* boson transverse momentum distribution in pp collisions at $\sqrt{s} = 7$ TeV with the ATLAS detector. *JHEP* **09**, 145 (2014). doi:[10.1007/JHEP09\(2014\)145](https://doi.org/10.1007/JHEP09(2014)145). arXiv:[1406.3660](https://arxiv.org/abs/1406.3660) [hep-ex]
37. J. Pumplin et al., New generation of parton distributions with uncertainties from global QCD analysis. *JHEP* **07**, 012 (2002). doi:[10.1088/1126-6708/2002/07/012](https://doi.org/10.1088/1126-6708/2002/07/012). arXiv:[hep-ph/0201195](https://arxiv.org/abs/hep-ph/0201195)
38. D.J. Lange, The EvtGen particle decay simulation package. *Nucl. Instrum. Methods A* **462**, 152 (2001). doi:[10.1016/S0168-9002\(01\)00089-4](https://doi.org/10.1016/S0168-9002(01)00089-4) [hep-ex]
39. T. Gleisberg et al., Event generation with SHERPA 1.1. *JHEP* **02**, 007 (2009). doi:[10.1088/1126-6708/2009/02/007](https://doi.org/10.1088/1126-6708/2009/02/007). arXiv:[0811.4622](https://arxiv.org/abs/0811.4622) [hep-ph]
40. S. Schumann, F. Krauss, A parton shower algorithm based on Catani–Seymour dipole factorisation. *JHEP* **03**, 038 (2008). doi:[10.1088/1126-6708/2008/03/038](https://doi.org/10.1088/1126-6708/2008/03/038). arXiv:[0709.1027](https://arxiv.org/abs/0709.1027) [hep-ph]
41. S. Höche, F. Krauss, S. Schumann, F. Siegert, QCD matrix elements and truncated showers. *JHEP* **05**, 053 (2009). doi:[10.1088/1126-6708/2009/05/053](https://doi.org/10.1088/1126-6708/2009/05/053). arXiv:[0903.1219](https://arxiv.org/abs/0903.1219) [hep-ph]
42. T. Gleisberg, S. Höche, Comix, a new matrix element generator. *JHEP* **12**, 039 (2008). doi:[10.1088/1126-6708/2008/12/039](https://doi.org/10.1088/1126-6708/2008/12/039). arXiv:[0808.3674](https://arxiv.org/abs/0808.3674) [hep-ph]
43. F. Cascioli, P. Maierhofer, S. Pozzorini, Scattering amplitudes with open loops. *Phys. Rev. Lett.* **108**, 111601 (2012). doi:[10.1103/PhysRevLett.108.111601](https://doi.org/10.1103/PhysRevLett.108.111601). arXiv:[1111.5206](https://arxiv.org/abs/1111.5206) [hep-ph]
44. S. Höche, F. Krauss, M. Schönherr, F. Siegert, QCD matrix elements + parton showers: the NLO case. *JHEP* **04**, 027 (2013). doi:[10.1007/JHEP04\(2013\)027](https://doi.org/10.1007/JHEP04(2013)027). arXiv:[1207.5030](https://arxiv.org/abs/1207.5030) [hep-ph]
45. ATLAS Collaboration, Monte Carlo generators for the production of a W or Z/γ^* boson in association with jets at ATLAS in run 2. ATLAS-PHYS-PUB-2016-003 (2016). <https://cds.cern.ch/record/2120133>
46. ATLAS Collaboration, Further ATLAS tunes of PYTHIA 6 and Pythia 8. ATLAS-PHYS-PUB-2011-014 (2011). <https://cds.cern.ch/record/1400677>
47. A. Martin, W. Stirling, R. Thorne, G. Watt, Parton distributions for the LHC. *Eur. Phys. J. C* **63**, 189 (2009). doi:[10.1140/epjc/s10052-009-1072-5](https://doi.org/10.1140/epjc/s10052-009-1072-5). arXiv:[0901.0002](https://arxiv.org/abs/0901.0002) [hep-ph]
48. ATLAS Collaboration, The ATLAS simulation infrastructure. *Eur. Phys. J. C* **70**, 823 (2010). doi:[10.1140/epjc/s10052-010-1429-9](https://doi.org/10.1140/epjc/s10052-010-1429-9). arXiv:[1005.4568](https://arxiv.org/abs/1005.4568) [physics.ins-det]
49. S. Agostinelli et al., GEANT4: a simulation toolkit. *Nucl. Instrum. Methods Phys. Res. Sect. A* **506**, 250 (2003). doi:[10.1016/S0168-9002\(03\)01368-8](https://doi.org/10.1016/S0168-9002(03)01368-8)
50. ATLAS Collaboration, Expected photon performance in the ATLAS experiment. ATLAS-PHYS-PUB-2011-007 (2011). <https://cds.cern.ch/record/1345329>

51. ATLAS Collaboration, Electron and photon energy calibration with the ATLAS detector using LHC Run 1 data. *Eur. Phys. J. C* **74**, 3071 (2014). doi:[10.1140/epjc/s10052-014-3071-4](https://doi.org/10.1140/epjc/s10052-014-3071-4). arXiv:[1407.5063](https://arxiv.org/abs/1407.5063) [hep-ex]
52. ATLAS Collaboration, Measurement of the photon identification efficiencies with the ATLAS detector using LHC run-1 data. *Eur. Phys. J. C* **76**, 666 (2016). doi:[10.1140/epjc/s10052-016-4507-9](https://doi.org/10.1140/epjc/s10052-016-4507-9)
53. ATLAS Collaboration, Muon reconstruction performance of the ATLAS detector in proton–proton collision data at 13 TeV. *Eur. Phys. J. C* **76**, 292 (2016). doi:[10.1140/epjc/s10052-016-4120-y](https://doi.org/10.1140/epjc/s10052-016-4120-y). arXiv:[1603.05598](https://arxiv.org/abs/1603.05598) [hep-hp]
54. M. Cacciari, G.P. Salam, G. Soyez, The anti- k_t jet clustering algorithm. *JHEP* **04**, 063 (2008). doi:[10.1088/1126-6708/2008/04/063](https://doi.org/10.1088/1126-6708/2008/04/063). arXiv:[0802.1189](https://arxiv.org/abs/0802.1189) [hep-ph]
55. ATLAS Collaboration, Topological cell clustering in the ATLAS calorimeters and its performance in LHC run 1 (2016). arXiv:[1603.02934](https://arxiv.org/abs/1603.02934) [hep-ex]
56. ATLAS Collaboration, Jet energy scale measurements and their systematic uncertainties in proton–proton collisions at $\sqrt{s} = 13$ TeV with the ATLAS detector (2017). arXiv:[1703.09665](https://arxiv.org/abs/1703.09665) [hep-ex]
57. ATLAS Collaboration, A measurement of the calorimeter response to single hadrons and determination of the jet energy scale uncertainty using LHC Run-1 pp-collision data with the ATLAS detector. *Eur. Phys. J. C* **77**, 26 (2017). doi:[10.1140/epjc/s10052-016-4580-0](https://doi.org/10.1140/epjc/s10052-016-4580-0). arXiv:[1607.08842](https://arxiv.org/abs/1607.08842) [hep-ex]
58. ATLAS Collaboration, Performance of pile-up mitigation techniques for jets in pp collisions at $\sqrt{s} = 8$ TeV using the ATLAS detector. *Eur. Phys. J. C* **76**, 581 (2016). doi:[10.1140/epjc/s10052-016-4395-z](https://doi.org/10.1140/epjc/s10052-016-4395-z). arXiv:[1510.03823](https://arxiv.org/abs/1510.03823) [hep-ex]
59. ATLAS Collaboration, Performance of missing transverse momentum reconstruction in proton–proton collisions at 7 TeV with ATLAS. *Eur. Phys. J. C* **72**, 1844 (2012). doi:[10.1140/epjc/s10052-011-1844-6](https://doi.org/10.1140/epjc/s10052-011-1844-6). arXiv:[1108.5602](https://arxiv.org/abs/1108.5602) [hep-ex]
60. ATLAS Collaboration, Vertex reconstruction performance of the ATLAS detector at $\sqrt{s} = 13$ TeV. ATL-PHYS-PUB-2015-026 (2015). <https://cds.cern.ch/record/2037717>
61. ATLAS Collaboration, Performance of algorithms that reconstruct missing transverse momentum in $\sqrt{s} = 8$ TeV proton–proton collisions in the ATLAS detector (2016). arXiv:[1609.09324](https://arxiv.org/abs/1609.09324) [hep-ex]
62. ATLAS Collaboration, Measurement of the inclusive isolated prompt photon cross section in pp collisions at $\sqrt{s} = 7$ TeV with the ATLAS detector. *Phys. Rev. D* **83**, 052005 (2011). doi:[10.1103/PhysRevD.83.052005](https://doi.org/10.1103/PhysRevD.83.052005). arXiv:[1012.4389](https://arxiv.org/abs/1012.4389) [hep-ex]
63. ATLAS Collaboration, Luminosity determination in pp collisions at $\sqrt{s} = 8$ TeV using the ATLAS detector at the LHC. *Eur. Phys. J. C* **76**, 653 (2016). doi:[10.1140/epjc/s10052-016-4466-1](https://doi.org/10.1140/epjc/s10052-016-4466-1). arXiv:[1608.03953](https://arxiv.org/abs/1608.03953) [hep-ex]
64. ATLAS Collaboration, Selection of jets produced in proton–proton collisions with the ATLAS detector using 2011 data. ATLAS-CONF-2012-020 (2012). <https://cds.cern.ch/record/1430034>
65. L. Lonnblad, Correcting the color dipole cascade model with fixed order matrix elements. *JHEP* **05**, 046 (2002). doi:[10.1088/1126-6708/2002/05/046](https://doi.org/10.1088/1126-6708/2002/05/046). arXiv:[hep-ph/0112284](https://arxiv.org/abs/hep-ph/0112284)
66. ATLAS Collaboration, Improved luminosity determination in pp collisions at $\sqrt{s} = 7$ TeV using the ATLAS detector at the LHC. *Eur. Phys. J. C* **73**, 2518 (2013). doi:[10.1140/epjc/s10052-013-2518-3](https://doi.org/10.1140/epjc/s10052-013-2518-3). arXiv:[1302.4393v2](https://arxiv.org/abs/1302.4393v2) [hep-ex]
67. J. Butterworth et al., PDF4LHC recommendations for LHC run II. *J. Phys. G: Nucl. Part. Phys.* **43**, 023001 (2016). doi:[10.1088/0954-3899/43/2/023001](https://doi.org/10.1088/0954-3899/43/2/023001). arXiv:[1510.03865](https://arxiv.org/abs/1510.03865) [hep-ph]
68. A. Buckley et al., LHAPDF6: parton density access in the LHC precision era (2014). doi:[10.1140/epjc/s10052-015-3318-8](https://doi.org/10.1140/epjc/s10052-015-3318-8). arXiv:[1412.7420](https://arxiv.org/abs/1412.7420) [hep-ph]
69. ATLAS Collaboration, Jet energy resolution in proton-proton collisions at $\sqrt{s} = 7$ TeV recorded in 2010 with the ATLAS detector. *Eur. Phys. J. C* **73**, 2306 (2013). doi:[10.1140/epjc/s10052-013-2306-0](https://doi.org/10.1140/epjc/s10052-013-2306-0). arXiv:[1210.6210](https://arxiv.org/abs/1210.6210) [hep-ex]
70. ATLAS Collaboration, Electron reconstruction and identification efficiency measurements with the ATLAS detector using the 2011 LHC proton–proton collision data. *Eur. Phys. J. C* **74**, 2941 (2014). doi:[10.1140/epjc/s10052-014-2941-0](https://doi.org/10.1140/epjc/s10052-014-2941-0). arXiv:[1404.2240](https://arxiv.org/abs/1404.2240) [hep-ex]
71. A.L. Read, Presentation of search results: the CL(s) technique. *J. Phys. G* **28**, 2693 (2002). doi:[10.1088/0954-3899/28/10/313](https://doi.org/10.1088/0954-3899/28/10/313)
72. T. Junk et al., Confidence level computation for combining searches with small statistics. *Nucl. Instrum. Methods Phys. Res. Sect. A* **434**, 435 (1999). doi:[10.1016/S0168-9002\(99\)00498-2](https://doi.org/10.1016/S0168-9002(99)00498-2). arXiv:[hep-ex/9902006](https://arxiv.org/abs/hep-ex/9902006)
73. G. Cowan, K. Cranmer, E. Gross, O. Vitells, Asymptotic formulae for likelihood-based tests of new physics. *Eur. Phys. J. C* **71**, 1554 (2011). doi:[10.1140/epjc/s10052-011-1554-0](https://doi.org/10.1140/epjc/s10052-011-1554-0). arXiv:[1007.1727](https://arxiv.org/abs/1007.1727) [physics.data-an] [Erratum: *Eur. Phys. J. C* **73**, 2501 (2013)]
74. A. Albert et al., Recommendations of the LHC Dark Matter Working Group: comparing LHC searches for heavy mediators of dark matter production in visible and invisible decay channels (2017). arXiv:[1703.05703](https://arxiv.org/abs/1703.05703) [hep-ex]
75. F. Kahlhoefer et al., Implications of unitarity and gauge invariance for simplified dark matter models. *JHEP* **02**, 016 (2016). doi:[10.1007/JHEP02\(2016\)016](https://doi.org/10.1007/JHEP02(2016)016). arXiv:[1510.02110](https://arxiv.org/abs/1510.02110) [hep-ph]
76. PLANCK Collaboration, P.A.R. Ade et al., Planck 2015 results. XIII. Cosmological parameters. *Astron. Astrophys.* **594**, A13 (2016). doi:[10.1051/0004-6361/201525830](https://doi.org/10.1051/0004-6361/201525830). arXiv:[1502.01589](https://arxiv.org/abs/1502.01589) [astro-ph.CO]
77. A. Boveia et al., Recommendations on presenting LHC searches for missing transverse energy signals using simplified s -channel models of dark matter (2016). arXiv:[1603.04156](https://arxiv.org/abs/1603.04156) [hep-ex]
78. PICO60 Collaboration, C. Amole et al., Dark matter search results from the PICO-60 C₃F₈ bubble chamber (2017). arXiv:[1702.07666](https://arxiv.org/abs/1702.07666) [astro-ph.CO]
79. LUX Collaboration, D.S. Akerib et al., Results on the spin-dependent scattering of weakly interacting massive particles on nucleons from the run 3 data of the LUX experiment. *Phys. Rev. Lett.* **116**, 161302 (2016). doi:[10.1103/PhysRevLett.116.161302](https://doi.org/10.1103/PhysRevLett.116.161302). arXiv:[1602.03489](https://arxiv.org/abs/1602.03489) [hep-ex]
80. PandaX-II Collaboration, A. Tan et al., Dark matter results from first 98.7 days of data from the PandaX-II experiment. *Phys. Rev. Lett.* **117**, 121303 (2016). doi:[10.1103/PhysRevLett.117.121303](https://doi.org/10.1103/PhysRevLett.117.121303). arXiv:[1607.07400](https://arxiv.org/abs/1607.07400) [hep-ex]
81. LUX Collaboration, D.S. Akerib et al., Results from a search for dark matter in the complete LUX exposure. *Phys. Rev. Lett.* **118**, 021303 (2017). doi:[10.1103/PhysRevLett.118.021303](https://doi.org/10.1103/PhysRevLett.118.021303). arXiv:[1608.07648](https://arxiv.org/abs/1608.07648) [astro-ph.CO]
82. SuperCDMS Collaboration, R. Agnese et al., New results from the search for low-mass weakly interacting massive particles with the CDMS low ionization threshold experiment. *Phys. Rev. Lett.* **116**, 071301 (2016). doi:[10.1103/PhysRevLett.116.071301](https://doi.org/10.1103/PhysRevLett.116.071301). arXiv:[1509.02448](https://arxiv.org/abs/1509.02448) [astro-ph.CO]
83. CRESST Collaboration, G. Angloher, et al., Results on light dark matter particles with a low-threshold CRESST-II detector. *Eur. Phys. J. C* **76**, 25 (2016). doi:[10.1140/epjc/s10052-016-3877-3](https://doi.org/10.1140/epjc/s10052-016-3877-3). arXiv:[1509.01515](https://arxiv.org/abs/1509.01515) [astro-ph.CO]
84. ATLAS Collaboration, Search for heavy resonances decaying to a Z boson and a photon in pp collisions at $\sqrt{s} = 13$ TeV with the ATLAS detector. *Phys. Lett. B* **764**, 11–30 (2017). doi:[10.1016/j.physletb.2016.11.005](https://doi.org/10.1016/j.physletb.2016.11.005). arXiv:[1607.06363](https://arxiv.org/abs/1607.06363) [hep-ex]
85. ATLAS Collaboration, ATLAS computing acknowledgements 2016–2017. ATL-GEN-PUB-2016-002 (2016). <https://cds.cern.ch/record/2202407>

ATLAS Collaboration

M. Aaboud^{137d}, G. Aad⁸⁸, B. Abbott¹¹⁵, J. Abdallah⁸, O. Abdinov^{12,*}, B. Abeloos¹¹⁹, S. H. Abidi¹⁶¹, O. S. AbouZeid¹³⁹, N. L. Abraham¹⁵¹, H. Abramowicz¹⁵⁵, H. Abreu¹⁵⁴, R. Abreu¹¹⁸, Y. Abulaiti^{148a,148b}, B. S. Acharya^{167a,167b,a}, S. Adachi¹⁵⁷, L. Adamczyk^{41a}, J. Adelman¹¹⁰, M. Adersberger¹⁰², T. Adye¹³³, A. A. Affolder¹³⁹, T. Agatonovic-Jovin¹⁴, C. Agheorghiesei^{28c}, J. A. Aguilar-Saavedra^{128a,128f}, S. P. Ahlen²⁴, F. Ahmadov^{68,b}, G. Aielli^{135a,135b}, S. Akatsuka⁷¹, H. Akerstedt^{148a,148b}, T. P. A. Åkesson⁸⁴, A. V. Akimov⁹⁸, G. L. Alberghi^{22a,22b}, J. Albert¹⁷², P. Albicocco⁵⁰, M. J. Alconada Verzini⁷⁴, M. Aleksa³², I. N. Aleksandrov⁶⁸, C. Alexa^{28b}, G. Alexander¹⁵⁵, T. Alexopoulos¹⁰, M. Alhroob¹¹⁵, B. Ali¹³⁰, M. Aliev^{76a,76b}, G. Alimonti^{94a}, J. Alison³³, S. P. Alkire³⁸, B. M. M. Allbrooke¹⁵¹, B. W. Allen¹¹⁸, P. P. Allport¹⁹, A. Aloisio^{106a,106b}, A. Alonso³⁹, F. Alonso⁷⁴, C. Alpigiani¹⁴⁰, A. A. Alshehri⁵⁶, M. Alstady⁸⁸, B. Alvarez Gonzalez³², D. Álvarez Piqueras¹⁷⁰, M. G. Alviggi^{106a,106b}, B. T. Amadio¹⁶, Y. Amaral Coutinho^{26a}, C. Amelung²⁵, D. Amidei⁹², S. P. Amor Dos Santos^{128a,128c}, A. Amorim^{128a,128b}, S. Amoroso³², G. Amundsen²⁵, C. Anastopoulos¹⁴¹, L. S. Ancu⁵², N. Andari¹⁹, T. Andeen¹¹, C. F. Anders^{60b}, J. K. Anders⁷⁷, K. J. Anderson³³, A. Andreazza^{94a,94b}, V. Andrei^{60a}, S. Angelidakis⁹, I. Angelozzi¹⁰⁹, A. Angerami³⁸, A. V. Anisenkov^{111,c}, N. Anjos¹³, A. Annovi^{126a,126b}, C. Antel^{60a}, M. Antonelli⁵⁰, A. Antonov^{100,*}, D. J. Antrim¹⁶⁶, F. Anulli^{134a}, M. Aoki⁶⁹, L. Aperio Bella³², G. Arabidze⁹³, Y. Arai⁶⁹, J. P. Araque^{128a}, V. Araujo Ferraz^{26a}, A. T. H. Arce⁴⁸, R. E. Ardell⁸⁰, F. A. Arduh⁷⁴, J.-F. Arguin⁹⁷, S. Argyropoulos⁶⁶, M. Arik^{20a}, A. J. Armbruster¹⁴⁵, L. J. Armitage⁷⁹, O. Arnaez¹⁶¹, H. Arnold⁵¹, M. Arratia³⁰, O. Arslan²³, A. Artamonov⁹⁹, G. Artoni¹²², S. Artz⁸⁶, S. Asai¹⁵⁷, N. Asbah⁴⁵, A. Ashkenazi¹⁵⁵, L. Asquith¹⁵¹, K. Assamagan²⁷, R. Astalos^{146a}, M. Atkinson¹⁶⁹, N. B. Atlay¹⁴³, L. Aubry⁵⁸, K. Augsten¹³⁰, G. Avolio³², B. Axen¹⁶, M. K. Ayoub¹¹⁹, G. Azuelos^{97,d}, A. E. Baas^{60a}, M. J. Baca¹⁹, H. Bachacou¹³⁸, K. Bachas^{76a,76b}, M. Backes¹²², M. Backhaus³², P. Bagnaia^{134a,134b}, H. Bahrasemani¹⁴⁴, J. T. Baines¹³³, M. Bajic³⁹, O. K. Baker¹⁷⁹, E. M. Baldwin^{111,c}, P. Balek¹⁷⁵, F. Balli¹³⁸, W. K. Balunas¹²⁴, E. Banas⁴², Sw. Banerjee^{176,e}, A. A. E. Bannoura¹⁷⁸, L. Barak³², E. L. Barberio⁹¹, D. Barberis^{53a,53b}, M. Barbero⁸⁸, T. Barillari¹⁰³, M.-S. Barisits³², T. Barklow¹⁴⁵, N. Barlow³⁰, S. L. Barnes^{36c}, B. M. Barnett¹³³, R. M. Barnett¹⁶, Z. Barnovska-Blenessy^{36a}, A. Baroncelli^{136a}, G. Barone²⁵, A. J. Barr¹²², L. Barranco Navarro¹⁷⁰, F. Barreiro⁸⁵, J. Barreiro Guimarães da Costa^{35a}, R. Bartoldus¹⁴⁵, A. E. Barton⁷⁵, P. Bartos^{146a}, A. Basalae¹²⁵, A. Bassalat^{119,f}, R. L. Bates⁵⁶, S. J. Batista¹⁶¹, J. R. Batley³⁰, M. Battaglia¹³⁹, M. Bauce^{134a,134b}, F. Bauer¹³⁸, H. S. Bawa^{145,g}, J. B. Beacham¹¹³, M. D. Beattie⁷⁵, T. Beau⁸³, P. H. Beauchemin¹⁶⁵, P. Bechtel²³, H. P. Beck^{18,h}, K. Becker¹²², M. Becker⁸⁶, M. Beckingham¹⁷³, C. Becot¹¹², A. J. Beddall^{20d}, A. Beddall^{20b}, V. A. Bednyakov⁶⁸, M. Bedognetti¹⁰⁹, C. P. Bee¹⁵⁰, T. A. Beermann³², M. Begalli^{26a}, M. Begel²⁷, J. K. Behr⁴⁵, A. S. Bell⁸¹, G. Bella¹⁵⁵, L. Bellagamba^{22a}, A. Bellerive³¹, M. Bellomo¹⁵⁴, K. Belotskiy¹⁰⁰, O. Beltramello³², N. L. Belyaev¹⁰⁰, O. Benary^{155,*}, D. Bencheekroun^{137a}, M. Bender¹⁰², K. Bendtz^{148a,148b}, N. Benekos¹⁰, Y. Benhammou¹⁵⁵, E. Benhar Nocchioli¹⁷⁹, J. Benitez⁶⁶, D. P. Benjamin⁴⁸, M. Benoit⁵², J. R. Bensinger²⁵, S. Bentvelsen¹⁰⁹, L. Beresford¹²², M. Beretta⁵⁰, D. Berge¹⁰⁹, E. Bergeas Kuutmann¹⁶⁸, N. Berger⁵, J. Beringer¹⁶, S. Berlendis⁵⁸, N. R. Bernard⁸⁹, G. Bernardi⁸³, C. Bernius¹⁴⁵, F. U. Bernlochner²³, T. Berry⁸⁰, P. Berta¹³¹, C. Bertella^{35a}, G. Bertoli^{148a,148b}, F. Bertolucci^{126a,126b}, I. A. Bertram⁷⁵, C. Bertsche⁴⁵, D. Bertsche¹¹⁵, G. J. Besjes³⁹, O. Bessidskaia Bylund^{148a,148b}, M. Bessner⁴⁵, N. Besson¹³⁸, C. Betancourt⁵¹, A. Bethani⁸⁷, S. Bethke¹⁰³, A. J. Bevan⁷⁹, J. Beyer¹⁰³, R. M. Bianchi¹²⁷, O. Biebel¹⁰², D. Biedermann¹⁷, R. Bielski⁸⁷, N. V. Biesuz^{126a,126b}, M. Biglietti^{136a}, J. Bilbao De Mendizabal⁵², T. R. V. Billoud⁹⁷, H. Bilokon⁵⁰, M. Bindi⁵⁷, A. Bingul^{20b}, C. Bini^{134a,134b}, S. Biondi^{22a,22b}, T. Bisanz⁵⁷, C. Bittrich⁴⁷, D. M. Bjergaard⁴⁸, C. W. Black¹⁵², J. E. Black¹⁴⁵, K. M. Black²⁴, R. E. Blair⁶, T. Blazek^{146a}, I. Bloch⁴⁵, C. Blocker²⁵, A. Blue⁵⁶, W. Blum^{86,*}, U. Blumenschein⁷⁹, S. Blunier^{34a}, G. J. Bobbink¹⁰⁹, V. S. Bobrovnikov^{111,c}, S. S. Bocchetta⁸⁴, A. Bocchi⁴⁸, C. Bock¹⁰², M. Boehler⁵¹, D. Boerner¹⁷⁸, D. Bogavac¹⁰², A. G. Bogdanchikov¹¹¹, C. Bohm^{148a}, V. Boisvert⁸⁰, P. Bokan^{168,i}, T. Bold^{41a}, A. S. Boldyrev¹⁰¹, A. E. Bolz^{60b}, M. Bomben⁸³, M. Bona⁷⁹, M. Boonekamp¹³⁸, A. Borisov¹³², G. Borissov⁷⁵, J. Bortfeldt³², D. Bortoletto¹²², V. Bortolotto^{62a,62b,62c}, D. Boscherini^{22a}, M. Bosman¹³, J. D. Bossio Sola²⁹, J. Boudreau¹²⁷, J. Bouffard², E. V. Bouhova-Thacker⁷⁵, D. Boumediene³⁷, C. Bourdarios¹¹⁹, S. K. Boutle⁵⁶, A. Boveia¹¹³, J. Boyd³², I. R. Boyko⁶⁸, J. Bracinik¹⁹, A. Brandt⁸, G. Brandt⁵⁷, O. Brandt^{60a}, U. Bratzler¹⁵⁸, B. Brau⁸⁹, J. E. Brau¹¹⁸, W. D. Breaden Madden⁵⁶, K. Brendlinger⁴⁵, A. J. Brennan⁹¹, L. Brenner¹⁰⁹, R. Brenner¹⁶⁸, S. Bressler¹⁷⁵, D. L. Briglin¹⁹, T. M. Bristow⁴⁹, D. Britton⁵⁶, D. Britzger⁴⁵, F. M. Brochu³⁰, I. Brock²³, R. Brock⁹³, G. Brooijmans³⁸, T. Brooks⁸⁰, W. K. Brooks^{34b}, J. Brosamer¹⁶, E. Brost¹¹⁰, J. H. Broughton¹⁹, P. A. Bruckman de Renstrom⁴², D. Bruncko^{146b}, A. Bruni^{22a}, G. Bruni^{22a}, L. S. Bruni¹⁰⁹, B. H. Brunt³⁰, M. Bruschi^{22a}, N. Bruscino²³, P. Bryant³³, L. Bryngemark⁴⁵, T. Buanes¹⁵, Q. Buat¹⁴⁴, P. Buchholz¹⁴³, A. G. Buckley⁵⁶, I. A. Budagov⁶⁸, F. Buehrer⁵¹, M. K. Bugge¹²¹, O. Bulekov¹⁰⁰, D. Bullock⁸, T. J. Burch¹¹⁰, H. Burckhart³², S. Burdin⁷⁷, C. D. Burgard⁵¹, A. M. Burger⁵, B. Burghgrave¹¹⁰, K. Burka⁴², S. Burke¹³³, I. 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