Theory and Applications of VHF Meteor Radar Observations

Joel P. Younger

Thesis submitted for the degree of Doctor of Philosophy in

Physics

Supervisors

Prof. Iain M. Reid and Prof. Robert A. Vincent The University of Adelaide School of Chemistry and Physics

> submitted 25 October, 2011

Contents

\mathbf{A}	Abstract					
Declaration						
Acknowledgments						
1	Intr	Introduction				
	1.1	Meteo	rs	10		
	1.2	The A	tmosphere	12		
	1.3	Meteo	r Observations	12		
	1.4	Early	History of Meteor Radar	14		
	1.5	Thesis	Structure	16		
	1.6	Summ	ary of Original Research	18		
2	Interferometric Meteor Radar			21		
	2.1	1 Hardware Configuration				
		2.1.1	Antenna Array	21		
		2.1.2	ATRAD Meteor Radar General Configuration	22		
		2.1.3	Transmission Signal Flow	26		
		2.1.4	Reception Signal Flow	27		
		2.1.5	Phase Determination	28		
	2.2	2.2 Angle-of-Arrival Estimation				
		2.2.1	0.5/2.0/4.5 method	29		

		2.2.2	Conversion to azimuth, zenith, and height $\ldots \ldots \ldots \ldots$	31
	2.3	Perfor	mance	32
		2.3.1	Underdense meteor function	32
		2.3.2	Overdense cutoff	34
		2.3.3	Minimum detectable electron line density	35
		2.3.4	Observed detection rates	35
2.4 Antenna Mutual Coupling		na Mutual Coupling	37	
		2.4.1	Measurement of Mutual Coupling	38
		2.4.2	Impact of Mutual Coupling on Positional Accuracy	40
	2.5	Calibr	ation	42
		2.5.1	Meteor echo phase calibration	42
		2.5.2	Correction using geophysical data	43
		2.5.3	Equi- D shell fitting \ldots \ldots \ldots \ldots \ldots \ldots \ldots	44
		2.5.4	Effectiveness of equi- D shell fitting $\ldots \ldots \ldots \ldots \ldots \ldots$	45
	2.6	Summ	ary	48
3	2.6 Atn	Summ nosphe	ary	48 51
3	2.6 Atn 3.1	Summ nosphe Meteo	ary	48 51 52
3	2.6 Atm 3.1 3.2	Summ nosphe Metec Therm	ary	 48 51 52 53
3	 2.6 Atm 3.1 3.2 3.3 	Summ nosphe Metec Thern Sputte	ary	48 51 52 53 56
3	 2.6 Atm 3.1 3.2 3.3 	Summ nosphe Metec Thern Sputte 3.3.1	ary	48 51 52 53 56 57
3	 2.6 Atm 3.1 3.2 3.3 	Summ nosphe Metec Therm Sputte 3.3.1 3.3.2	ary	48 51 52 53 56 57 57
3	 2.6 Atm 3.1 3.2 3.3 3.4 	Summ nosphe Metec Therm Sputte 3.3.1 3.3.2 Ioniza	ary	48 51 52 53 56 57 57 59
3	 2.6 Atm 3.1 3.2 3.3 3.4 3.5 	Summ nosphe Metec Thern Sputte 3.3.1 3.3.2 Ioniza Nume	ary	48 51 52 53 56 57 57 59 60
3	 2.6 Atm 3.1 3.2 3.3 3.4 3.5 	Summ nosphe Metec Therm Sputte 3.3.1 3.3.2 Ioniza Nume 3.5.1	ary	48 51 52 53 56 57 57 59 60 60
3	 2.6 Atm 3.1 3.2 3.3 3.4 3.5 	Summ nosphe Metec Thern Sputte 3.3.1 3.3.2 Ioniza Nume 3.5.1 3.5.2	ary	48 51 52 53 56 57 57 59 60 60 60 61
3	 2.6 Atm 3.1 3.2 3.3 3.4 3.5 	Summ nosphe Metec Thern Sputte 3.3.1 3.3.2 Ioniza Nume 3.5.1 3.5.2 3.5.3	ary	48 51 52 53 56 57 57 59 60 60 60 61 63
3	 2.6 Atm 3.1 3.2 3.3 3.4 3.5 	Summ nosphe Metec Therm Sputte 3.3.1 3.3.2 Ioniza Nume 3.5.1 3.5.2 3.5.3 3.5.4	ary	48 51 52 53 56 57 57 59 60 60 61 63 64

		3.5.6 General ablative behavior			
		3.5.7 Deceleration $\ldots \ldots \ldots$			
		3.5.8 Significance of sputtering			
	3.6	Summary			
4	Meteor Detection Heights				
	4.1	Numerical simulation			
	4.2	Effect of density scale height			
		4.2.1 Linearity of H^* vs $h_1 - h_2$			
	4.3	Effect of constant density level			
	4.4	Observations			
	4.5	Summary			
-	D				
9	Pro	perties of Underdense Meteor Echoes 95			
	5.1	Fresnel Diffraction Pattern			
	5.2	Trail Evolution			
		5.2.1 Initial radius $\ldots \ldots 100$			
		5.2.2 Diffusion			
	5.3	Echo Decay			
	5.4	Effects of Trail Parameters on Echo Decay			
		5.4.1 Effect of non-Gaussian initial distribution			
		5.4.2 Non-constant electron line density			
		5.4.3 Finite velocity $\ldots \ldots \ldots$			
		5.4.4 Non-constant diffusion coefficient			
		5.4.5 Fragmentation			
	5.5	Velocity Estimation			
		5.5.1 Echo slope			
		5.5.2 Fresnel transform			
	5.6	Summary			

6	Anomalous Diffusion			117		
	6.1	Multip	ble Ionic Species	. 121		
		6.1.1	Numerical simulation method	. 121		
		6.1.2	Simulation results: common initial radius	. 123		
		6.1.3	Simulation results: individual initial radii	. 125		
		6.1.4	General considerations	. 129		
	6.2	Aeroso	ol Absorption	. 130		
		6.2.1	Numerical simulation method	. 131		
		6.2.2	Numerical simulation predictions	. 132		
	6.3	Compa	arison with observations	. 137		
	6.4	Summ	ary	. 139		
7	Met	Meteor Shower Detection 143				
	7.1	Great	circle method	. 145		
		7.1.1	Properties of the acceptance band	. 147		
	7.2	2006-2	2007 Meteor shower survey	. 149		
		7.2.1	Velocity estimation	. 151		
		7.2.2	Detected shower radiants	. 151		
		7.2.3	Uncertainty in the velocity estimate	. 156		
	7.3	Stream	n orbits	. 158		
		7.3.1	Orbital elements	. 158		
		7.3.2	Infall effects	. 160		
		7.3.3	Equations of orbital motion	. 161		
		7.3.4	Orbits of detected shower streams	. 162		
		7.3.5	Kozai resonance	. 166		
	7.4	Summ	ary	. 167		
8	Conclusions 169					
	8.1	Meteo	r Radar Performance	. 169		
	8.2	Meteo	roid Ablation	. 170		

	8.3	Meteor Detection Height	•	171		
	8.4	Underdense Echoes	•	172		
	8.5	Anomalous Diffusion Phenomena	•	174		
	8.6	Meteor Astronomy	•	175		
	8.7	Future Work	•	176		
Α	Glo	ssary of Mathematical Terms]	179		
в	B Published Paper: Geophys. Res. Lett., 35(L15812):1-4, 2008					
\mathbf{C}	Puł	olished Paper: Mon. Not. R. astr. Soc., 398: 350-356, 2009]	191		

v

vi

Abstract

This thesis examines the operation and observations made by VHF interferometric radar. Broad topics include the operation of interferometric meteor radar, the physics of meteor ablation, the formation and diffusion of meteor trails, and meteor astronomy.

The performance of the basic radar configuration is examined with particular attention paid to the source and mitigation of positional errors introduced by hardware. Sources include random errors in phase and range estimates, mutual coupling between antennas, and biases in the phase measurements used to determine the angle-of-arrival of incident radiation. A new method for post-statistical steering is presented, using height dependent ambipolar diffusion coefficients as a reference.

The physics of meteor flight and ablation in the atmosphere is examined in detail. The heating and vaporization of meteoroids as they enter Earth's atmosphere is modeled and the effect of sputtering on the formation of meteor ionization is assessed. These calculations are performed for a variety of meteor types, sorted by velocity, angle-of-entry, composition, and size. The results are then used in conjunction with atmospheric models to produce predictions of meteor radar performance.

The effect of the atmosphere on the ablation and subsequent detection of meteors is considered and used to construct new metrics for the characterization of the atmospheric density profile in the meteor region. The effects of constant density level and density scale height are assessed with regards to the peak detection height and range of heights over which a radar detects meteors.

The formation of the underdense meteor echo is examined in detail. New contributions to the understanding of this topic include an assessment of the effects of variable electron line density in the trail, deceleration, and fragmentation on the eventual measurement of the decay time of meteor echoes. Estimates of ambipolar diffusion coefficients are examined by determining the effect of anomalous diffusion resulting from electron absorbing aerosols and multi-constituent trail chemistry.

Meteor astronomy techniques used to overcome the limitations of interferometric meteor radars are implemented in order to search for discrete streams of solar system debris that result in meteor showers. The results include a significant number of previously undiscovered shower systems. I certify that this work contains no material which has been accepted for the award of any degree or diploma in any university or tertiary institution and, to the best of my knowledge, contains no material previously published or written by another person, except where due reference has been made in the text.

I give consent to this copy of my thesis, when deposited in the University of Adelaide Library, being made available for loan and photocopying, subject to the provisions of the Copyright Act of 1968.

I also give permission for the digital version of my thesis to be made available on the internet, via the University of Adelaide digital research repository, the University of Adelaide Library catalogue, the Australasian Digital Theses Program (ADTP) and through internet search engines.

Signed date

Joel P. Younger, BSc (Hons.)

Acknowledgments

The process that takes a graduate student from application to the submission of a thesis is one that cannot occur without the generous help and professional, emotional, and financial support of a huge number of people. I would be remiss if I did not take the opportunity to thank as many of them as possible, although I suspect that I will inevitably miss some in this brief note.

My supervisors during the last three and a half years, Professor Iain Reid and Professor Robert Vincent have provided the support, guidance, and freedom that I feel has enabled me to mature as a scientist, rather than simply accomplish the feat of writing a thesis. They have given me a degree of independence in my research that less confident supervisors might balk at, allowing me to focus on new and productive lines of inquiry. With their encouragement to undertake a PhD in the first place, careful contributions when I became lost, and tireless efforts to assist me in establishing a place in the broader scientific community, it has been a pleasure and an invaluable learning experience to be under their tutelage. Good bosses are rare. I have had two.

Perhaps most of all, my partner Josie deserves credit for the successful completion of this thesis, for putting up with me playing grad student when by all rights I should have been out getting a real job. She has put up with the absence and absent-mindedness that are inherent to the scientifically minded. To our beautiful son Miles, you make the simple act of walking through the door an event. Whatever you grow up to be I hope that you discover some of the joy of finding things out that I have come to know. In the mean time, you can have as many bananas as you want.

Without the support and the values instilled by my family, I would not have even

considered nor been able to begin down this path. My father Stephen and my mother Mari did an exceptional job in teaching me to find something I enjoy and work hard at it. They have always supported me in my education and otherwise, and for that I cannot thank them enough. My brother James has always been a steadfast friend, as comfortable in inane banter after a year away as after an hour away.

It is perhaps a sign of how lucky I am to be surrounded by good people that I find the usual stigma associated with in-laws to be so foreign. While I have been busying myself with research and conferences and thesis writing, they have supplied a steady stream of dinners, barbecues, and child minding. To Dianne and Michael, Frank and Julia, Marina, Todd, and Oli, Bianca and Nathan, Nonna, the wider Barbaro clan, and the Hibberts, thank you for all the help you have provided to Josie, Miles and myself. You may not realize it, but you have been an essential component of this research.

Over the course of the work described in this thesis, it has been my pleasure to work with some exceptional scientists. Damian Murphy of the Australian Antarctic Division has been generous beyond reason with data, which made much of the work in this thesis possible in the first place. David Holdsworth has provided the answers to many of my questions on the obscura of meteor radar, and his work in the field provides a foundation for much of my own.

The Atmospheric Physics Research group at the University of Adelaide also deserves some mention (or perhaps blame?). Like any good graduate student environment, it has been a rotating crew of motley figures, all good company and offering something to learn, a strange blend of calculus, pretty graphs, and beer. Dr. Andrew MacKinnon stands out as perhaps one of the most helpful postdocs in the history of professional science. He has had the answers to so many questions on the minutia of radar design and operation, it has been like having a personal help-line just down the hall. Of course, there are also my co-conspirators in grad school life, Dr. Peter Love, Dr. Bronwyn Dolman, and Ray Oermann, as well as special guest appearances by Octavianus Cakra Satya, Dr. Sujata Kovalam, and Dr. Jens Lautenbach.

A special mention must be made of the legendary Dr. Graham Elford, who has been

kind enough to periodically venture out of retirement to assist me with comments and conversation. Few grad students are fortunate enough to be able to talk with someone who has been so integral to their field of research, let alone someone as approachable and helpful as Dr. Elford.

I have been lucky to be just a short drive away from the manufacturer of most of my equipment, ATRAD Pty. Ltd. They have been more than accommodating in my requests for their time and resources, as well as making some fine radars when grad students are not pestering them. Chris Adami showed profound patience as I tried to grasp the functions of basic components and Richard Mayo pretended to not notice when I nearly broke the brand new vectore impedance meter that he drove out to BP for me to fool around with.

This research would not have been possible without significant financial support. The University of Adelaide School of Chemistry and Physics and the Australian Research Council provided a generous scholarship that enabled me to take the time to complete this work. Much of the actual research was made possible by ASAC grant 2529 and ARC grants DP0878144 and DP0558361. I would also like to thank the Australian Institute of Physics for the STSP Student Paper prize and the organizers of the MST 12 conference for the Student Award.

In addition to the people I can mention by name, I would like to thank the anonymous reviewers of my published work. Your unattributed efforts have helped me to grow as a scientist and your comments and suggestions have provided valuable contributions to my research.

A number of my former colleagues from the United States Navy are perhaps unaware of the contribution they made towards the completion of this research. It was the imparted skills of systematic analysis and the exposure to so many different types of sensors that inspired me to study physics in the first place, as well beginning my fascination with radar. The experience of working hard, over long hours in difficult conditions, taught me lessons not available in any educational institution that have been invaluable over the years. I would therefore like to make special mention of my classmates and instructors from Recruit Training Center Great Lakes, Naval Aircrewman Candidate School, Aviation Rescue Swimmer School, AW 'A' School, Airborne Acoustic Mission Course, HSL-41, and SERE. It was my pleasure and honor to serve among the talented pilots, aircrewmen, and maintainers of HSL-37, especially Detachment 9 'Night Wizards' and Detachment 7 'War Dawgs', as well as the crews of the USS Lake Erie and USS Fletcher, and the special operators that spent the summer of 1999 in the NAG. I am still, to this day, learning from those five years. IYOYAS.

Finally, I would like to thank Kelley, for what is my favorite piece of academic prose: "The immediate and incredulous conclusion was that the rocket had penetrated a meteor trail, an event so unlikely as to be embarrassing to report." [Kelley et al., 1998]. Charged dust in the mesosphere is not typically a source of comedy, but I laugh every time I read that passage. Bravo, sir. Bravo.

If anyone has not been mentioned, or not mentioned by name, I apologize. There are so many that deserve my thanks.